

# **Uma formulação integral para as teorias de Yang-Mills**

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# Plan of the talk

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Integral formulation of Maxwell's theory

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**Generalized non-abelian Stokes theorem**

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Gauge and Integrable theories

# Maxwell's equations

## Maxwell's equations

$$\vec{\nabla} \cdot \vec{E} = \frac{\rho}{\varepsilon_0}$$

$$\vec{\nabla} \times \vec{B} - \mu_0 \varepsilon_0 \frac{\partial \vec{E}}{\partial t} = \mu_0 \vec{J}$$

$$\vec{\nabla} \cdot \vec{B} = 0$$

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$$E_i = F_{0i}$$

$$B_i = -\frac{1}{2c} \varepsilon_{ijk} F_{jk}$$

$$j^\mu \equiv \frac{1}{\varepsilon_0} \left( \rho, -\frac{1}{c} J^i \right)$$

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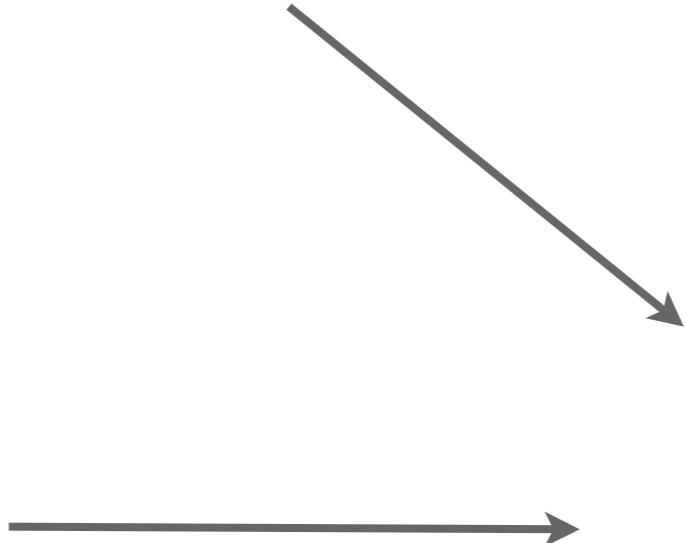
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$$\partial_\mu F^{\mu\nu} = j^\nu$$

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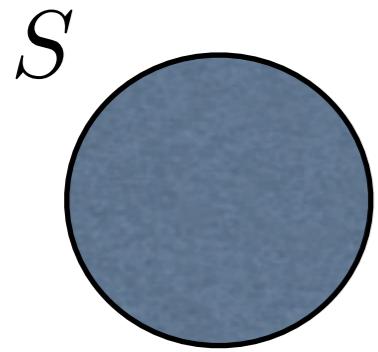
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$$\tilde{F}_{\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} F^{\rho\sigma}$$

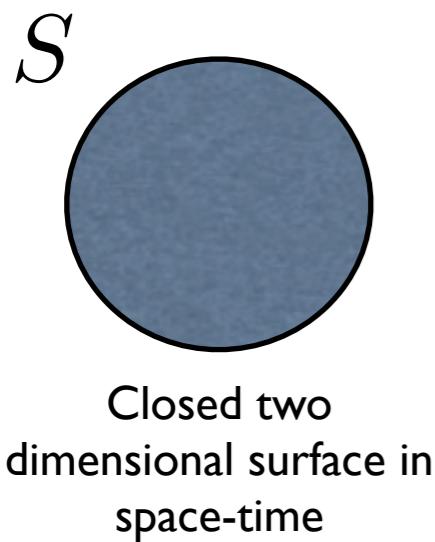
# Integral form of Maxwell's equations

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Closed two  
dimensional surface in  
space-time

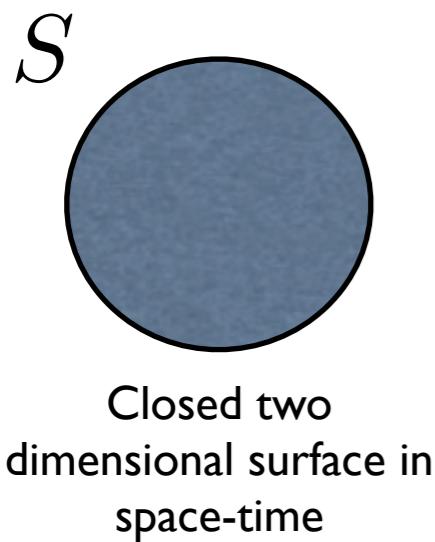
# Integral form of Maxwell's equations



$$\Phi(S) = \oint_S F_{\mu\nu} \frac{\partial x^\mu}{\partial \sigma} \frac{\partial x^\nu}{\partial \tau} d\sigma d\tau$$

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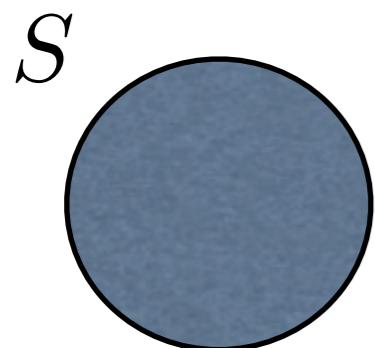
$$\Phi(S) = 0$$

**Maxwell's  
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$$\tilde{\Phi}(S) = -\frac{Q}{\epsilon_0}$$

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charge inside S

$$Q = -\epsilon_0 \int_V \tilde{J}_{\mu\nu\rho} \frac{\partial x^\mu}{\partial \sigma} \frac{\partial x^\nu}{\partial \tau} \frac{\partial x^\rho}{\partial \zeta} d\sigma d\tau d\zeta$$

$$\tilde{J}_{\mu\nu\rho} = \epsilon_{\mu\nu\rho\sigma} j^\sigma$$

$S$   $\equiv$  closed spatial surface

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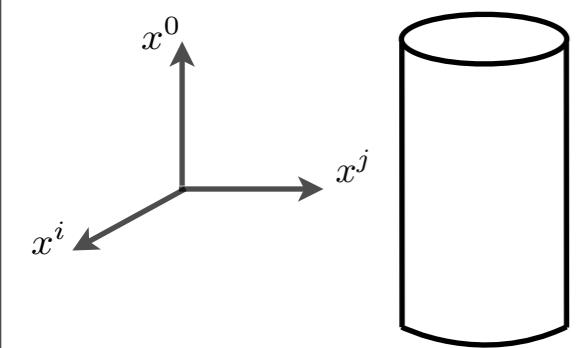
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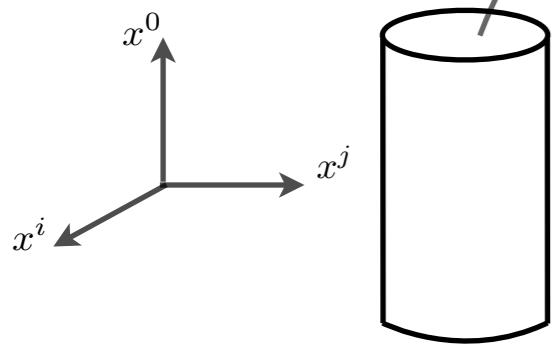


$S$  with a time component



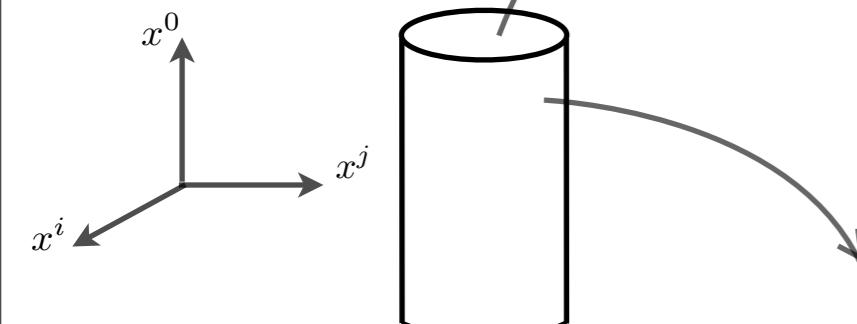
*S* with a time component

$$\Phi(\text{disc}) = \int_{\text{disc}} F_{ij} \frac{\partial x^i}{\partial \sigma} \frac{\partial x^j}{\partial \tau} d\sigma d\tau = -c \int_{\text{disc}} \vec{B} \cdot d\vec{\Sigma}$$



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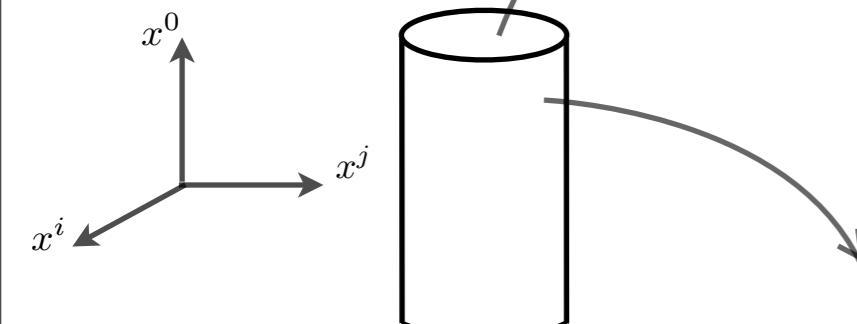
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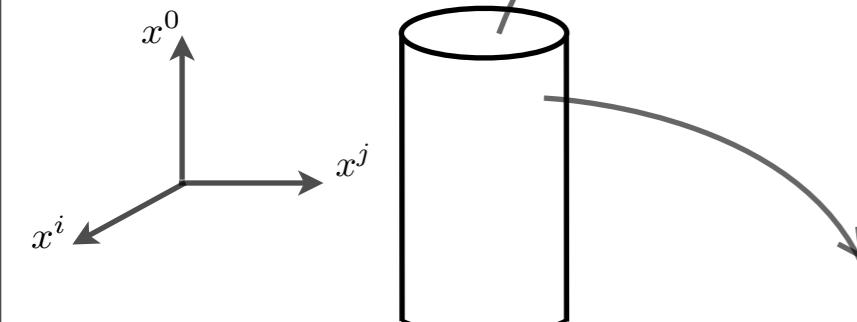
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Then

$$\Phi(S) = 0 = -c \int_{\text{top disc}} \vec{B} \cdot d\vec{\Sigma} + c \int_{\text{bottom disc}} \vec{B} \cdot d\vec{\Sigma} - \int dx^0 \oint \vec{E} \cdot d\vec{l}$$

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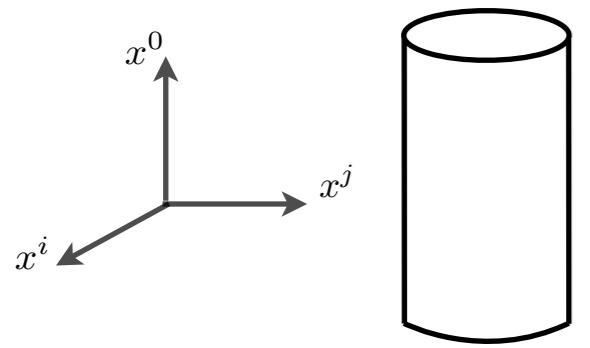
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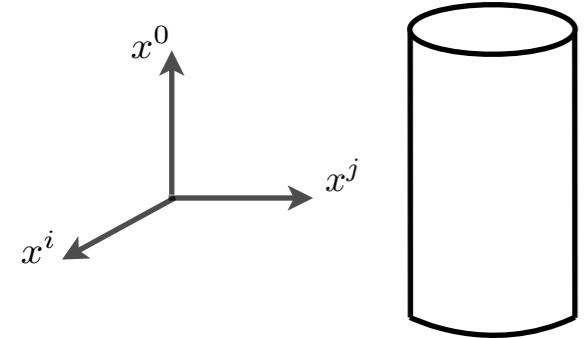
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In the limit  $\delta x^0 \rightarrow 0$

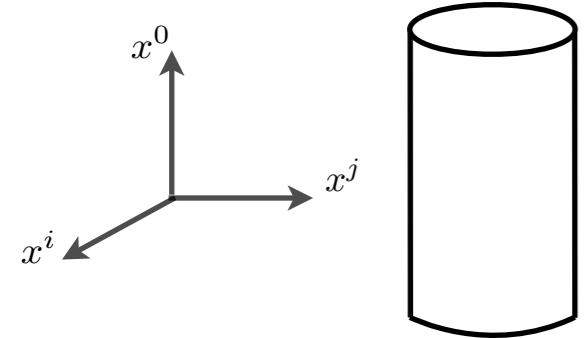
$$\frac{d}{dt} \int \vec{B} \cdot d\vec{\Sigma} = - \oint \vec{E} \cdot d\vec{l} \quad \rightarrow \quad \vec{\nabla} \times \vec{E} = - \frac{\partial \vec{B}}{\partial t}$$





Analogously

$$\begin{aligned}
 \tilde{\Phi}(S) &= - \int_{\text{top disc}} \vec{E} \cdot d\vec{\Sigma} + \int_{\text{bottom disc}} \vec{E} \cdot d\vec{\Sigma} + c \int dx^0 \oint \vec{B} \cdot d\vec{l} \\
 &= \frac{1}{c \varepsilon_0} \int dx^0 \int \vec{J} \cdot d\vec{\Sigma}
 \end{aligned}$$

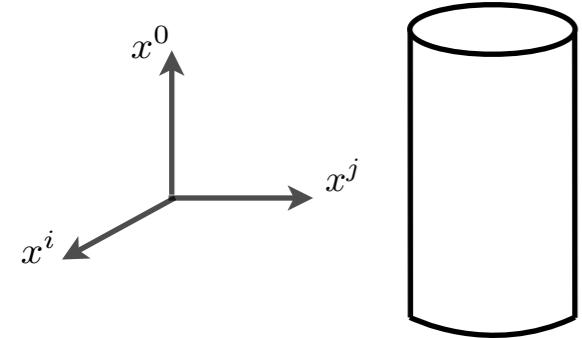


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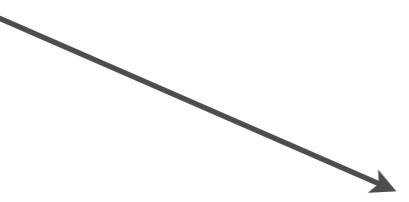


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# Summarizing

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Maxwell's eqs. are equivalent to

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V is the volume inside S

$$\left( \tilde{J}_{\mu\nu\rho} = \varepsilon_{\mu\nu\rho\sigma} j^\sigma \right)$$

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Maxwell's eqs. are recovered in the limit where S is infinitesimal

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Gauge covariant under the transformations

$$A_\mu \rightarrow g A_\mu g^{-1} + \frac{i}{e} \partial_\mu g g^{-1}$$

$$F_{\mu\nu} \rightarrow g F_{\mu\nu} g^{-1}$$

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## Charges

$$Q = \int d^3x \partial^i F_{i0} = \int d^3x \vec{\nabla} \cdot \vec{E} = \int d\vec{\Sigma} \cdot \vec{E}$$

$$\tilde{Q} = \int d^3x \partial^i \tilde{F}_{i0} = - \int d^3x \vec{\nabla} \cdot \vec{B} = - \int d\vec{\Sigma} \cdot \vec{B}$$

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$$Q \rightarrow \int d\vec{\Sigma} \cdot g \vec{E} g^{-1} = g Q g^{-1}$$

if  $g$  is constant at infinity

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## Under a gauge transformation

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eigenvalues of  $Q$   
are gauge invariant

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# The Quantum Theory of Fields

## Steven Weinberg

### vol. II, pag. 12-13

#### 15.3 Field Equations and Conservation Laws

Using Eq. (15.2.9) for the matrix  $g_{\alpha\beta}$  in Eq. (15.2.3), the full Lagrangian density is

$$\mathcal{L} = -\frac{1}{4} F_{\alpha\mu\nu} F_{\alpha}^{\mu\nu} + \mathcal{L}_M(\psi, D_\mu\psi), \quad (15.3.1)$$

where in the absence of gauge fields  $\mathcal{L}_M(\psi, \partial_\mu\psi)$  would be the ‘matter’ Lagrangian density. We could, in principle, include a dependence of  $\mathcal{L}_M$  on  $F_{\alpha\mu\nu}$  as well as higher covariant derivatives  $D_\nu D_\mu\psi$ ,  $D_\lambda F_{\alpha\mu\nu}$ , etc., but we exclude these non-renormalizable terms here for the same reason as in electrodynamics: as discussed in Section 12.3, such terms would be highly suppressed at ordinary energies by negative powers of some very large mass. For this reason the standard model of the weak, electromagnetic and strong interactions has a Lagrangian of the general form (15.3.1).

The equations of motion of the gauge field are here

$$\begin{aligned} \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_{\alpha\nu})} &= -\partial_\mu F_{\alpha}^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial A_{\alpha\nu}} \\ &= -F_{\gamma}^{\nu\mu} C_{\gamma\alpha\beta} A_{\beta\mu} - i \frac{\partial \mathcal{L}_M}{\partial D_\nu\psi} t_\alpha\psi \end{aligned}$$

and so

$$\partial_\mu F_{\alpha}^{\mu\nu} = -J_{\alpha}^\nu, \quad (15.3.2)$$

where  $J_{\alpha}^\nu$  is the current:

$$J_{\alpha}^\nu \equiv -F_{\gamma}^{\nu\mu} C_{\gamma\alpha\beta} A_{\beta\mu} - i \frac{\partial \mathcal{L}_M}{\partial D_\nu\psi} t_\alpha\psi. \quad (15.3.3)$$

The current  $J_{\alpha}^\nu$  is conserved in the ordinary sense

$$\partial_\nu J_{\alpha}^\nu = 0, \quad (15.3.4)$$

#### 15.3 Field Equations and Conservation Laws

13

as can be seen either from the Euler–Lagrange equations for  $\psi$  and the invariance equivalent (15.2.2) or, more easily, directly from the field equations (15.3.2).

The derivatives in Eqs. (15.3.2) and (15.3.4) are ordinary derivatives, not the gauge-covariant derivatives  $D_\nu$ , so the gauge invariance of these equations is somewhat obscure. It can be made manifest by rewriting Eq. (15.3.2) in terms of the gauge-covariant derivative of the field strength

$$\begin{aligned} D_\lambda F_{\alpha}^{\mu\nu} &\equiv \partial_\lambda F_{\alpha}^{\mu\nu} - i(t_\beta^A)_{\alpha\gamma} A_{\beta\lambda} F_{\gamma}^{\mu\nu} \\ &= \partial_\lambda F_{\alpha}^{\mu\nu} - C_{\alpha\gamma\beta} A_{\beta\lambda} F_{\gamma}^{\mu\nu}. \end{aligned} \quad (15.3.5)$$

Then Eq. (15.3.2) reads

$$D_\mu F_{\alpha}^{\mu\nu} = -J_{\alpha}^\nu, \quad (15.3.6)$$

where  $J_{\alpha}^\nu$  is the current of the matter fields alone

$$J_{\alpha}^\nu \equiv -i \frac{\partial \mathcal{L}_M}{\partial D_\nu\psi} t_\alpha\psi. \quad (15.3.7)$$

This is gauge-covariant, if  $\mathcal{L}_M$  is gauge-invariant. Also, by operating on Eq. (15.3.6) with  $D_\nu$ , using the commutation relation

$$[D_\nu, D_\mu] F_{\alpha}^{\rho\sigma} = -i(t_\gamma^A)_{\alpha\beta} F_{\gamma\nu\mu} F_{\beta}^{\rho\sigma} = -C_{\gamma\alpha\beta} F_{\gamma\nu\mu} F_{\beta}^{\rho\sigma},$$

we see that  $J_{\alpha}^\nu$  satisfies a gauge-covariant conservation law

$$D_\nu J_{\alpha}^\nu = 0, \quad (15.3.8)$$

rather than the ordinary conservation law (15.3.4) obeyed by the full current  $J_{\alpha}^\nu$ . Also, it is straightforward (using Eq. (15.1.5)) to derive the identities:

$$D_\mu F_{\alpha\nu\lambda} + D_\nu F_{\alpha\lambda\mu} + D_\lambda F_{\alpha\mu\nu} = 0, \quad (15.3.9)$$

which hold whether or not the gauge fields satisfy the field equations.

# The Quantum Theory of Fields

## Steven Weinberg

### vol. II, pag. 12-13

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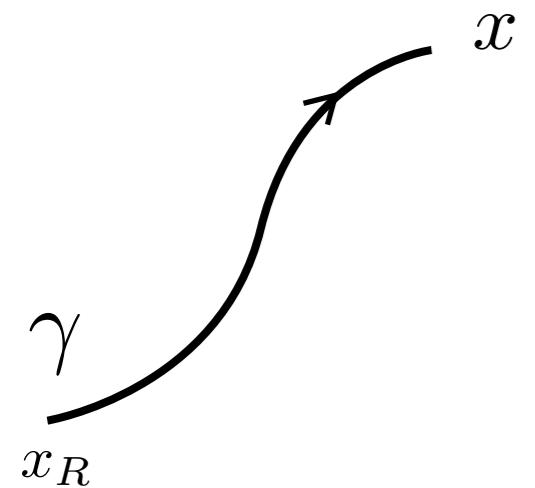
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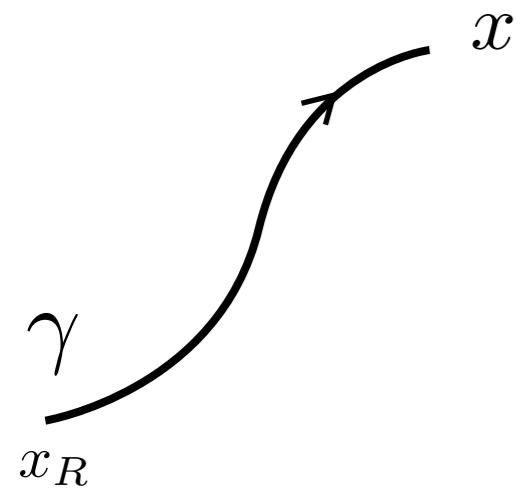
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# The Wilson line

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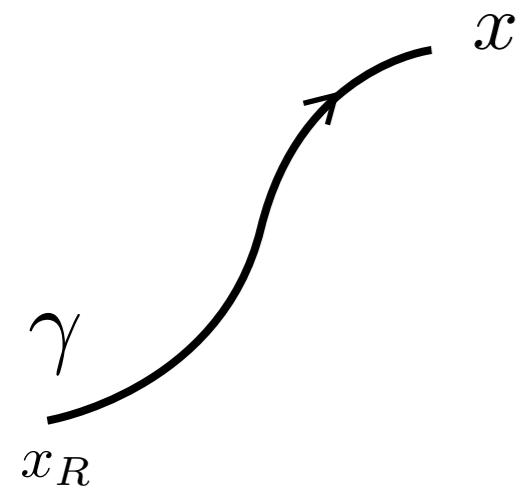


# The Wilson line



$$\frac{dW}{d\sigma} + i e A_\mu \frac{dx^\mu}{d\sigma} W = 0$$

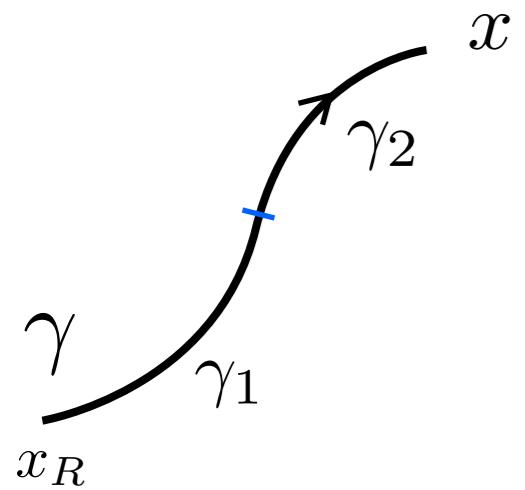
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$$\frac{dW}{d\sigma} + i e A_\mu \frac{dx^\mu}{d\sigma} W = 0$$

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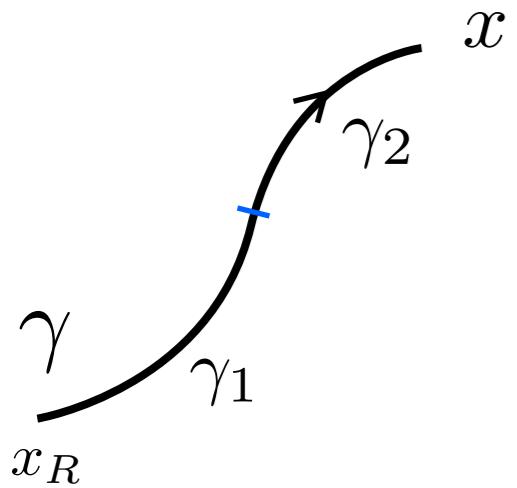


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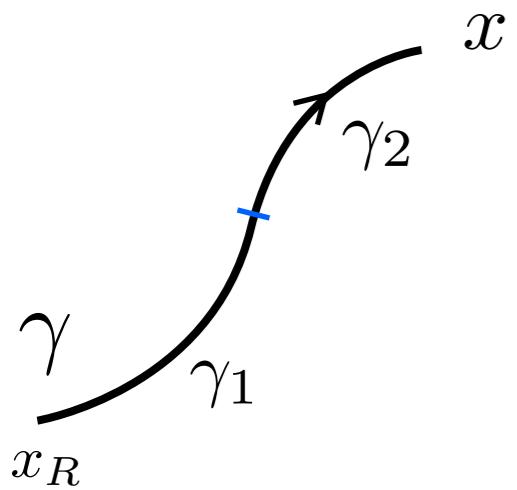
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2)  $A_\mu \rightarrow g A_\mu g^{-1} + \frac{i}{e} \partial_\mu g g^{-1}$       then       $W \rightarrow g(x) W g^{-1}(x_R)$

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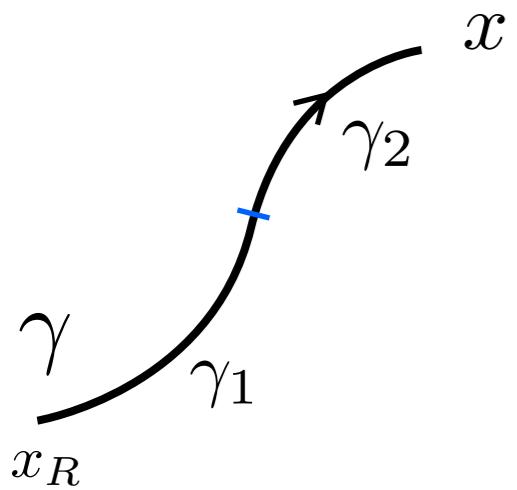
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3) Suppose  $B_{\mu\nu} \rightarrow g B_{\mu\nu} g^{-1}$

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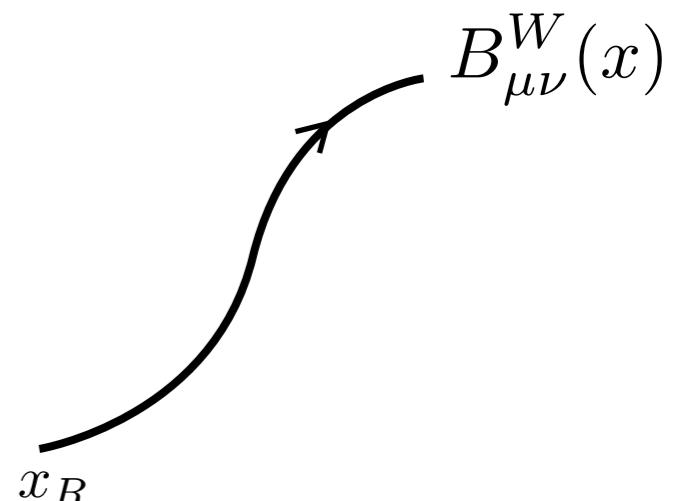
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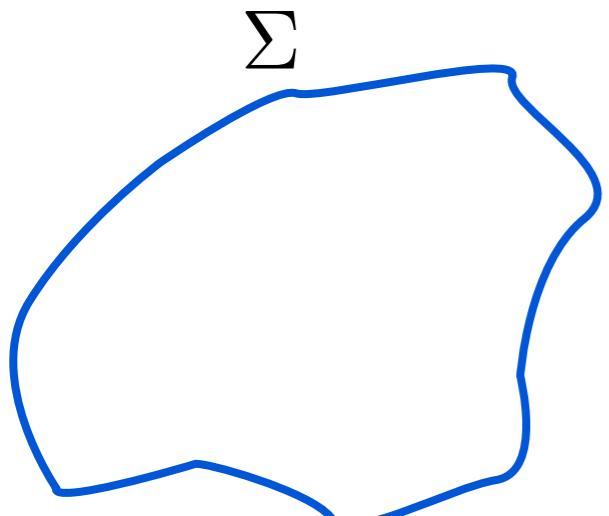
Then  $B_{\mu\nu}^W \equiv W^{-1} B_{\mu\nu} W$  satisfies

$$B_{\mu\nu}^W \rightarrow g(x_R) B_{\mu\nu}^W g^{-1}(x_R)$$

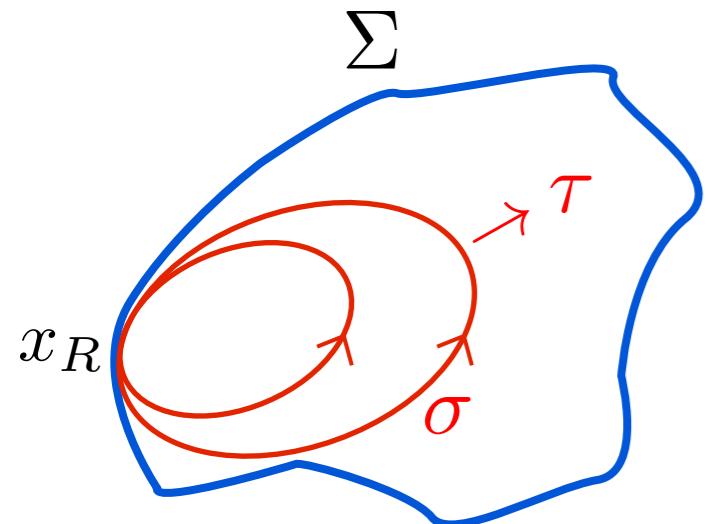


# Integrating on surfaces

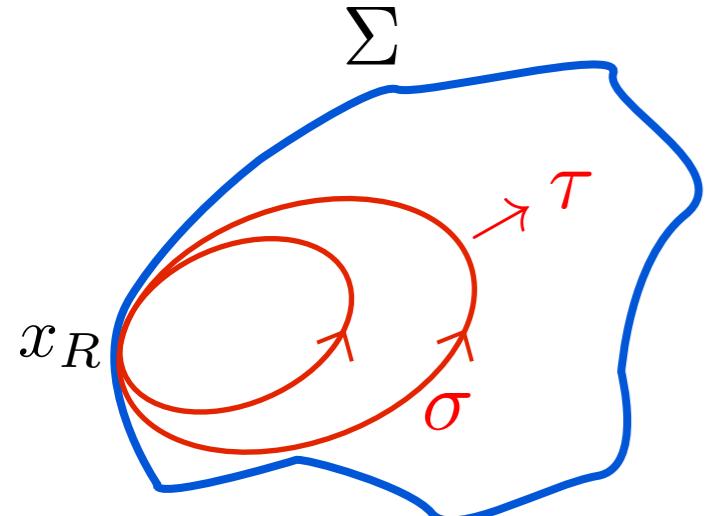
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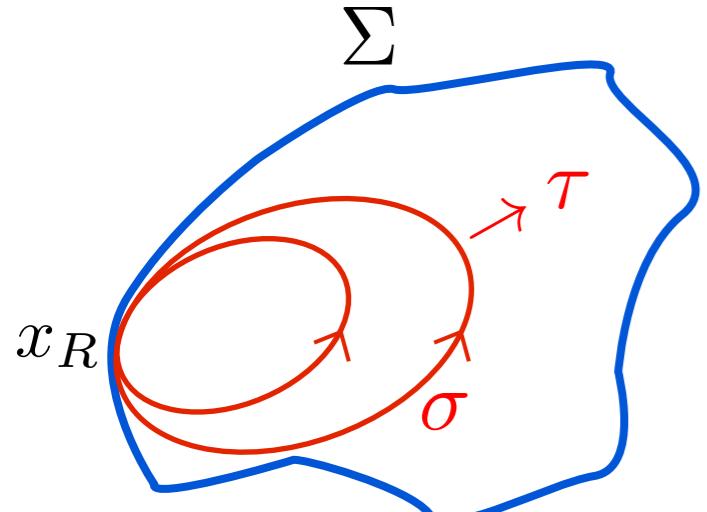
# Integrating on surfaces



$$\frac{dV}{d\tau} - VT(A, B, \tau) = 0$$

$$T(B, A, \tau) \equiv \int_0^{2\pi} d\sigma W^{-1} B_{\mu\nu} W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}$$

# Integrating on surfaces



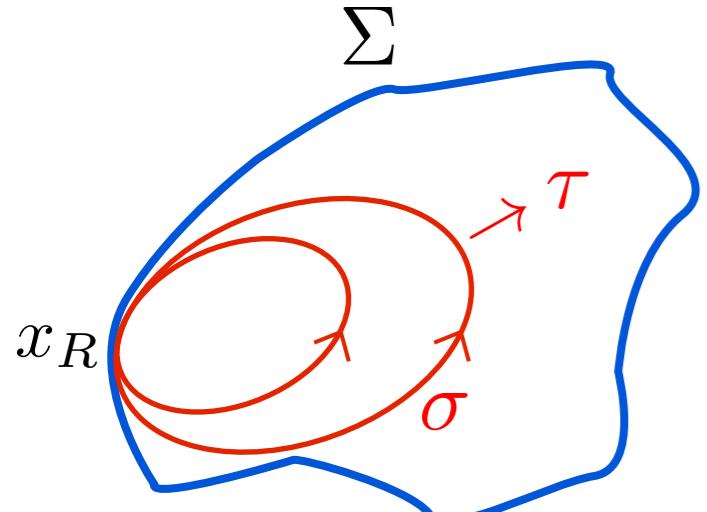
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It is a surface ordered integral

$$V(\Sigma) = V_R P_2 e^{\int_\Sigma d\sigma d\tau W^{-1} B_{\mu\nu} W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}}$$

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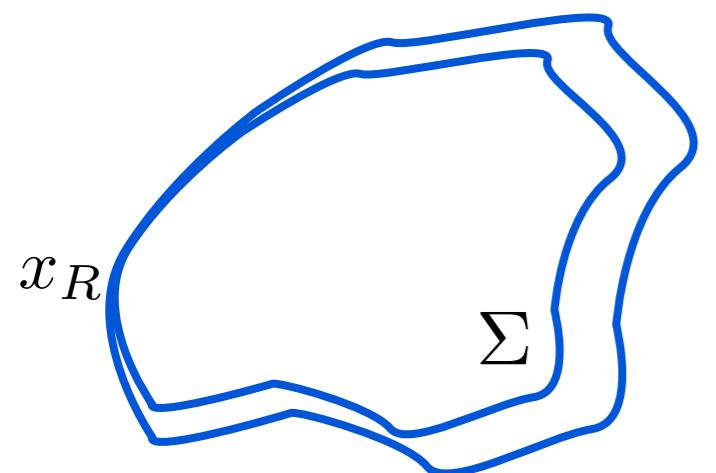
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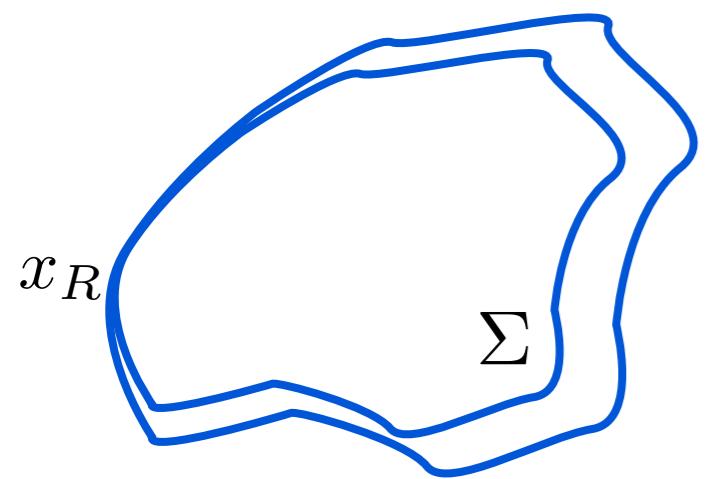
$$V(\Sigma) = V_R P_2 e^{\int_\Sigma d\sigma d\tau W^{-1} B_{\mu\nu} W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}}$$

$$\begin{aligned} A_\mu &\rightarrow g A_\mu g^{-1} + \frac{i}{e} \partial_\mu g g^{-1} \\ B_{\mu\nu} &\rightarrow g B_{\mu\nu} g^{-1} \end{aligned} \rightarrow \left\{ \begin{array}{l} T(B, A, \tau) \rightarrow g(x_R) T(B, A, \tau) g^{-1}(x_R) \\ V \rightarrow g(x_R) V g^{-1}(x_R) \end{array} \right.$$

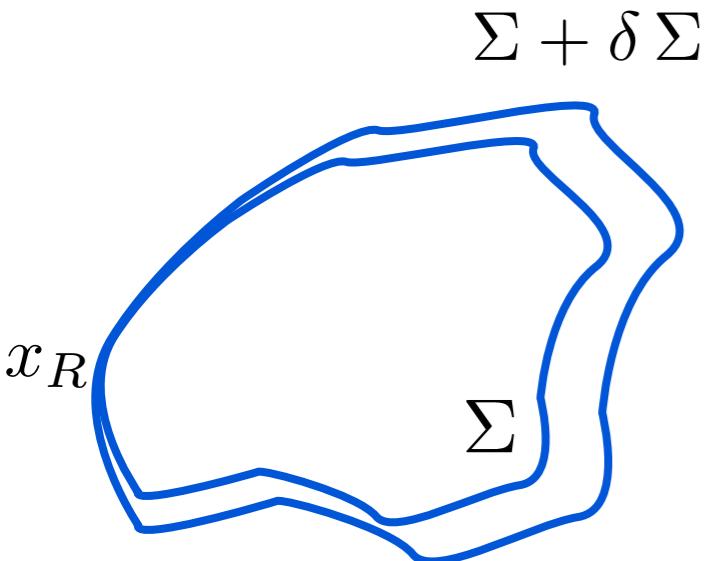


$\Sigma + \delta \Sigma$ 

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What happens to  $V$  when one changes the surface



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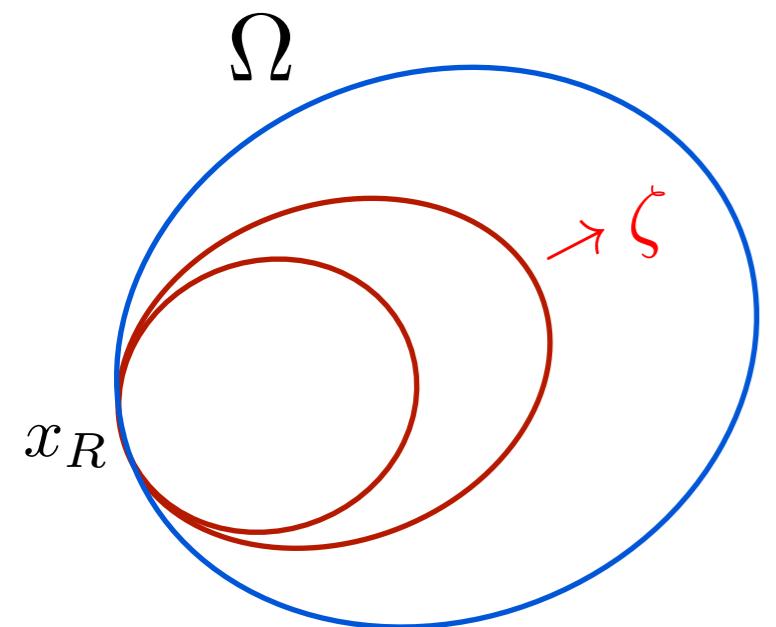
$$\begin{aligned}
 \delta V V^{-1} &\equiv \int_0^{2\pi} d\tau \int_0^{2\pi} d\sigma V(\tau) \{ \\
 &W^{-1} [D_\lambda B_{\mu\nu} + D_\mu B_{\nu\lambda} + D_\nu B_{\lambda\mu}] W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau} \delta x^\lambda \\
 &- \int_0^\sigma d\sigma' [B_{\kappa\rho}^W(\sigma') - ieF_{\kappa\rho}^W(\sigma'), B_{\mu\nu}^W(\sigma)] \frac{dx^\kappa}{d\sigma'} \frac{dx^\mu}{d\sigma} \\
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# Integrating on volumes

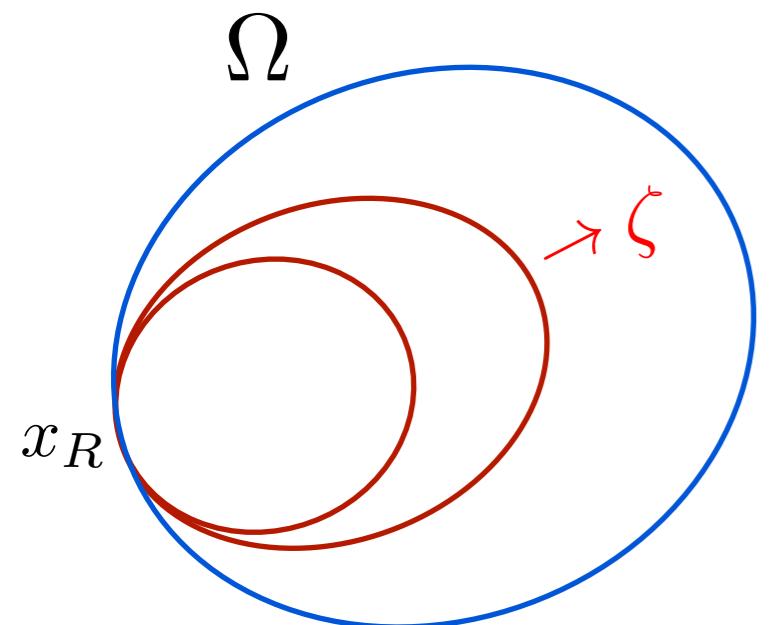
$\Omega$ 

# Integrating on volumes

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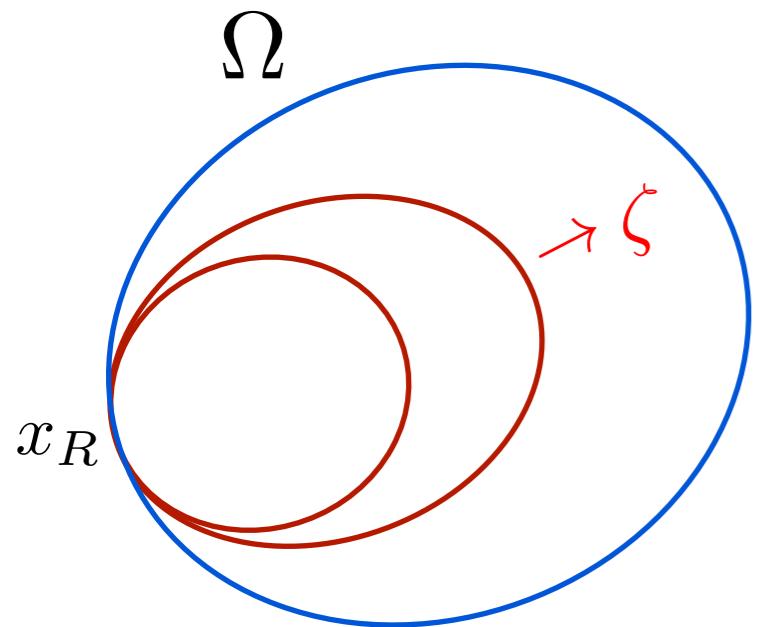


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$$\frac{dV}{d\zeta} - \mathcal{K}V = 0$$

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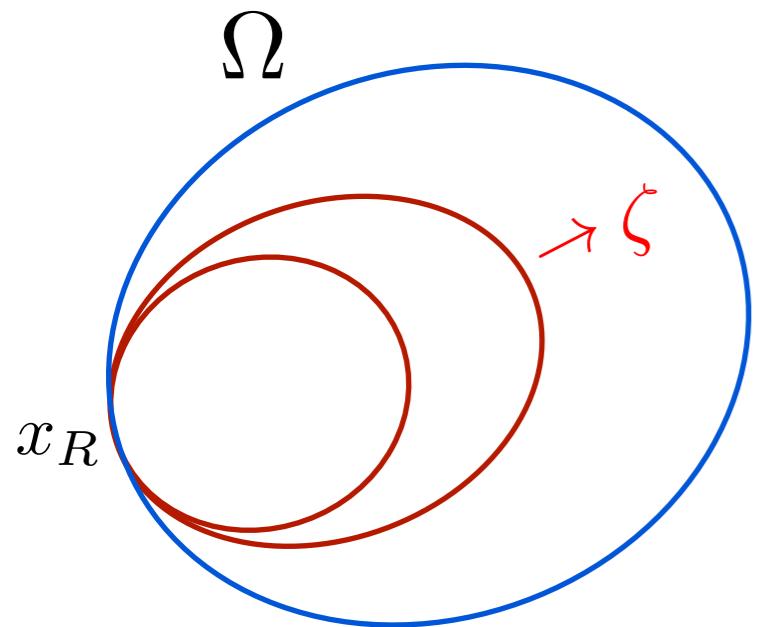


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$$W^{-1} [D_\lambda B_{\mu\nu} + D_\mu B_{\nu\lambda} + D_\nu B_{\lambda\mu}] W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau} \frac{dx^\lambda}{d\zeta} \\ - \int_0^\sigma d\sigma' \left[ B_{\kappa\rho}^W(\sigma') - ieF_{\kappa\rho}^W(\sigma'), B_{\mu\nu}^W(\sigma) \right] \frac{dx^\kappa}{d\sigma'} \frac{dx^\mu}{d\sigma} \\ \times \left( \frac{dx^\rho(\sigma')}{d\tau} \frac{dx^\nu(\sigma)}{d\zeta} - \frac{dx^\rho(\sigma')}{d\zeta} \frac{dx^\nu(\sigma)}{d\tau} \right) \} V^{-1}(\tau)$$

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Ordered volume integral →  $V(\partial\Omega) = P_3 e^{\int_\Omega \mathcal{K}} V_R$

# The generalized non-abelian Stokes theorem

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O. Alvarez, L. A. Ferreira and J. Sanchez Guillen,  
Nucl. Phys. B **529**, 689 (1998) [arXiv:hep-th/9710147].  
Int. J. Mod. Phys. A **24**, 1825 (2009) [arXiv:0901.1654 [hep-th]]

# The integral equations for Yang-Mills

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$$P_2 e^{ie \int_{\partial\Omega} d\tau d\sigma [\alpha F_{\mu\nu}^W + \beta \tilde{F}_{\mu\nu}^W] \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}} = P_3 e^{\int_\Omega d\zeta d\tau V \mathcal{J} V^{-1}}$$

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$$\begin{aligned} \mathcal{J} \equiv & \int_0^{2\pi} d\sigma \left\{ ie\beta \tilde{J}_{\mu\nu\lambda}^W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau} \frac{dx^\lambda}{d\zeta} + e^2 \int_0^\sigma d\sigma' \right. \\ & \times \left[ \left( (\alpha - 1) F_{\kappa\rho}^W + \beta \tilde{F}_{\kappa\rho}^W \right) (\sigma') , \left( \alpha F_{\mu\nu}^W + \beta \tilde{F}_{\mu\nu}^W \right) (\sigma) \right] \\ & \times \frac{dx^\kappa}{d\sigma'} \frac{dx^\mu}{d\sigma} \left. \left( \frac{dx^\rho(\sigma')}{d\tau} \frac{dx^\nu(\sigma)}{d\zeta} - \frac{dx^\rho(\sigma')}{d\zeta} \frac{dx^\nu(\sigma)}{d\tau} \right) \right\} \end{aligned}$$

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$$B_{\mu\nu} \rightarrow \alpha F_{\mu\nu} + \beta \tilde{F}_{\mu\nu}$$

# The integral equations for Yang-Mills

$$P_2 e^{ie \int_{\partial\Omega} d\tau d\sigma [\alpha F_{\mu\nu}^W + \beta \tilde{F}_{\mu\nu}^W] \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}} = P_3 e^{\int_\Omega d\zeta d\tau V \mathcal{J} V^{-1}}$$

$$\begin{aligned} \mathcal{J} \equiv & \int_0^{2\pi} d\sigma \left\{ ie\beta \tilde{J}_{\mu\nu\lambda}^W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau} \frac{dx^\lambda}{d\zeta} + e^2 \int_0^\sigma d\sigma' \right. \\ & \times \left[ \left( (\alpha - 1) F_{\kappa\rho}^W + \beta \tilde{F}_{\kappa\rho}^W \right) (\sigma') , \left( \alpha F_{\mu\nu}^W + \beta \tilde{F}_{\mu\nu}^W \right) (\sigma) \right] \\ & \times \frac{dx^\kappa}{d\sigma'} \frac{dx^\mu}{d\sigma} \left( \frac{dx^\rho(\sigma')}{d\tau} \frac{dx^\nu(\sigma)}{d\zeta} - \frac{dx^\rho(\sigma')}{d\zeta} \frac{dx^\nu(\sigma)}{d\tau} \right) \} \end{aligned}$$

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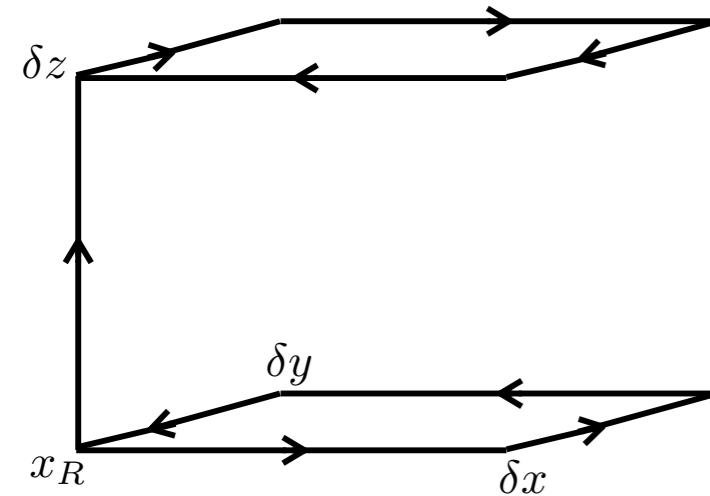
$$J^\mu = \frac{1}{3!} \varepsilon^{\mu\nu\rho\lambda} \tilde{J}_{\nu\rho\lambda}$$

direct consequence of Stokes theorem  
and Yang-Mills eqs.

It implies the local Yang-Mills equations

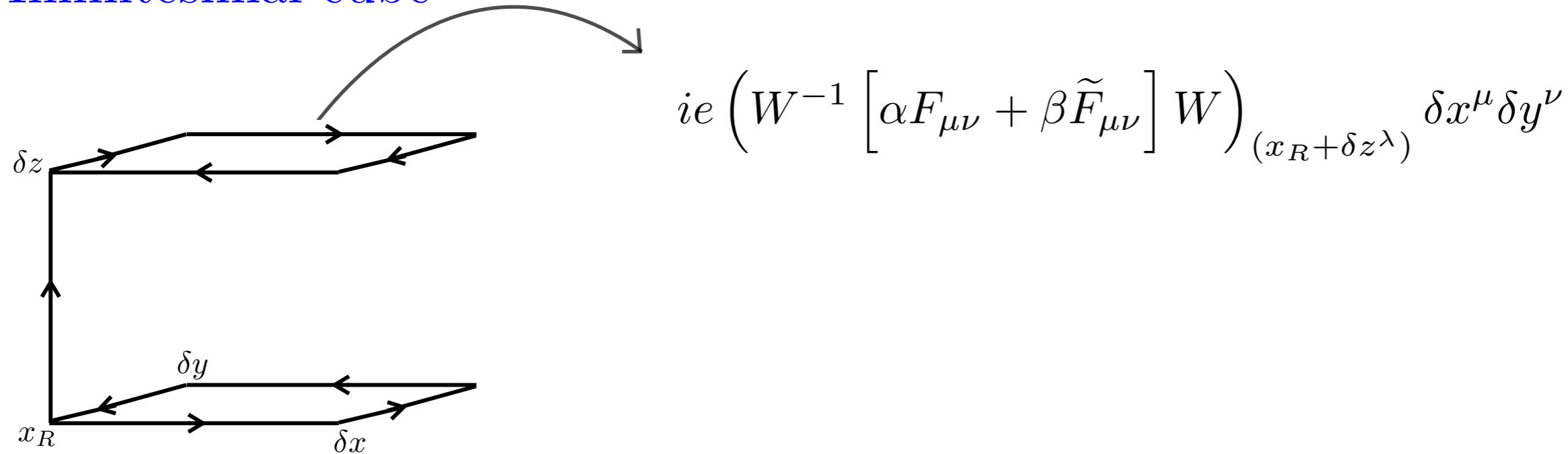
It implies the local Yang-Mills equations

Infinitesimal cube



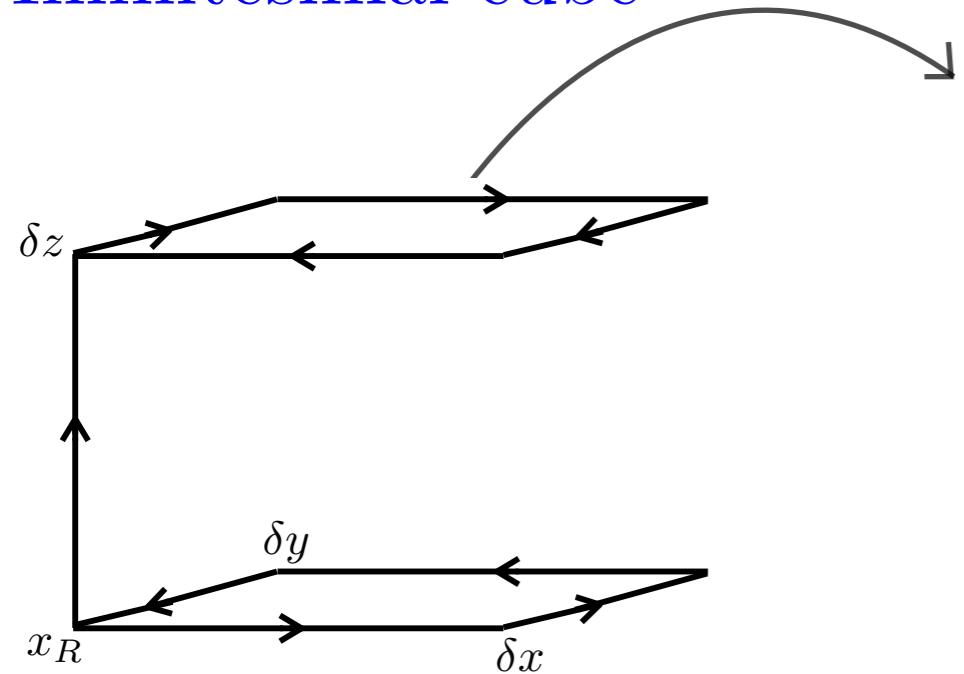
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# It implies the local Yang-Mills equations

Infinitesimal cube



$$ie \left( W^{-1} \left[ \alpha F_{\mu\nu} + \beta \tilde{F}_{\mu\nu} \right] W \right)_{(x_R + \delta z^\lambda)} \delta x^\mu \delta y^\nu$$

$$W_{(x_R + \delta z^\lambda)} \sim 1 - ie A_\lambda(x_R) \delta x^\lambda$$

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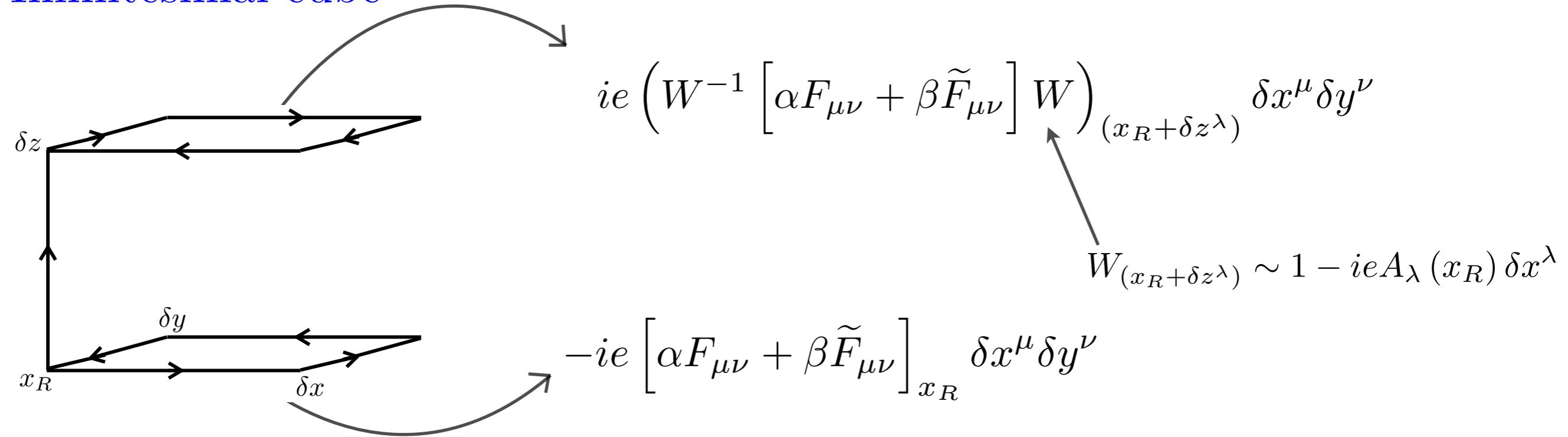
Infinitesimal cube

The diagram shows a small cube centered at  $x_R$ . The vertical edge is labeled  $\delta z$ , the horizontal edge along the front face is labeled  $\delta y$ , and the depth edge is labeled  $\delta x$ . Arrows indicate the direction of each edge.

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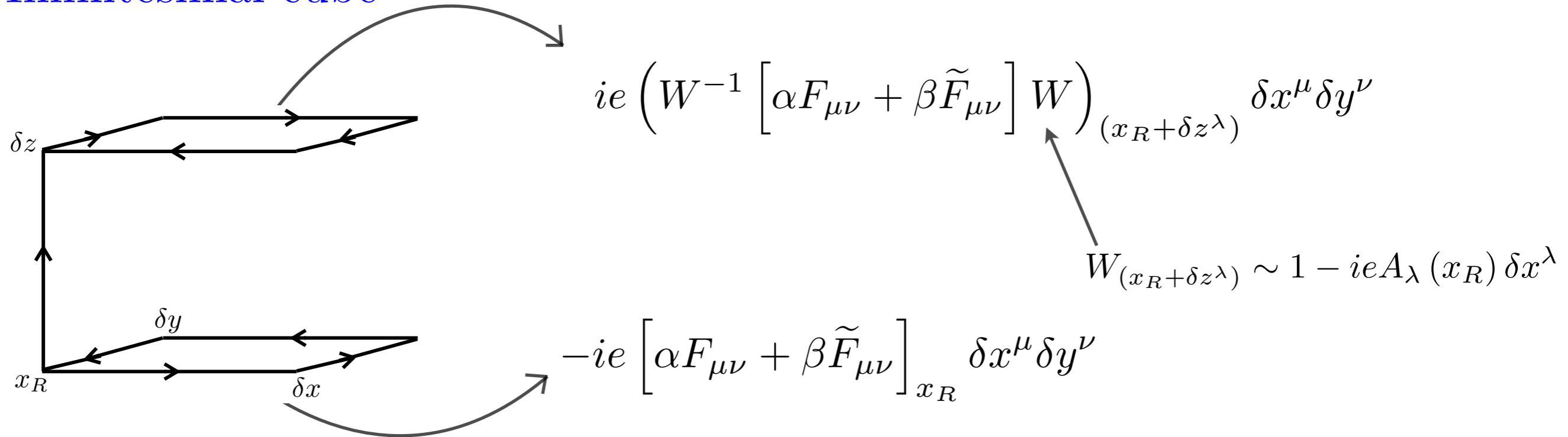


## Taylor expanding and adding other faces

$$P_2 e^{ie \int_{\partial\Omega} d\tau d\sigma [\alpha F_{\mu\nu}^W + \beta \tilde{F}_{\mu\nu}^W] \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}} \sim 1 + ie \left[ D_\lambda [\alpha F_{\mu\nu} + \beta \tilde{F}_{\mu\nu}]_{x_R} + \text{cyclic perm.} \right] \delta x^\mu \delta y^\nu \delta z^\lambda$$

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## The volume integral gives

$$P_3 e^{\int_{\Omega} d\zeta d\tau V \mathcal{J} V^{-1}} \sim 1 + ie \beta \tilde{J}_{\mu\nu\lambda} \delta x^\mu \delta y^\nu \delta z^\lambda$$

# Consequences

$$P_2 e^{ie \int_{\partial\Omega} d\tau d\sigma [\alpha F_{\mu\nu}^W + \beta \tilde{F}_{\mu\nu}^W] \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}} = P_3 e^{\int_\Omega d\zeta d\tau V \mathcal{J} V^{-1}}$$

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If  $\Omega_c$  is a closed volume ( $\partial\Omega_c = 0$ )

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|||  
$$q = \frac{2\pi n}{e\beta}$$

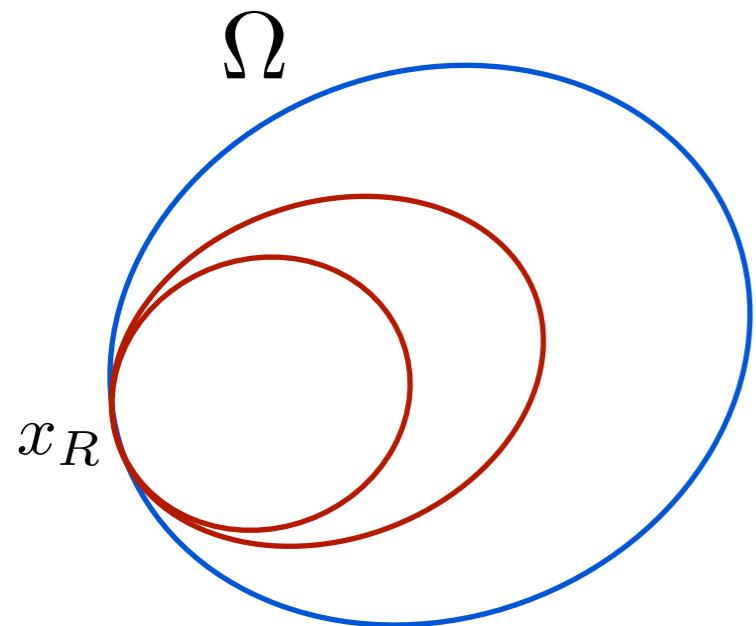
# Loop space

## Loop space

$$L\Omega = \{\gamma : S^2 \rightarrow \Omega \mid \text{north pole} \rightarrow x_R \in \partial\Omega\}$$

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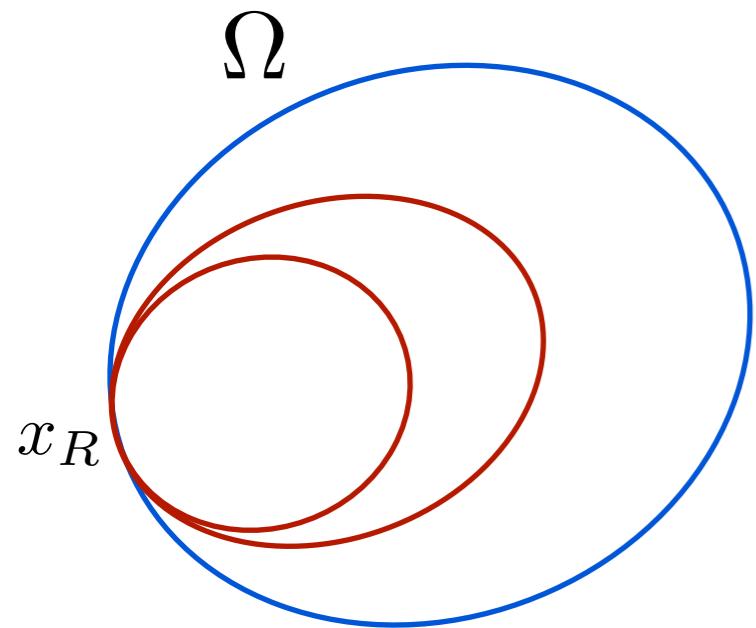
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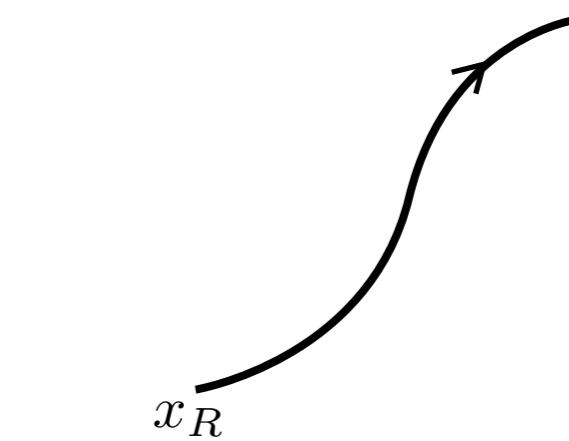
Volume in space-time

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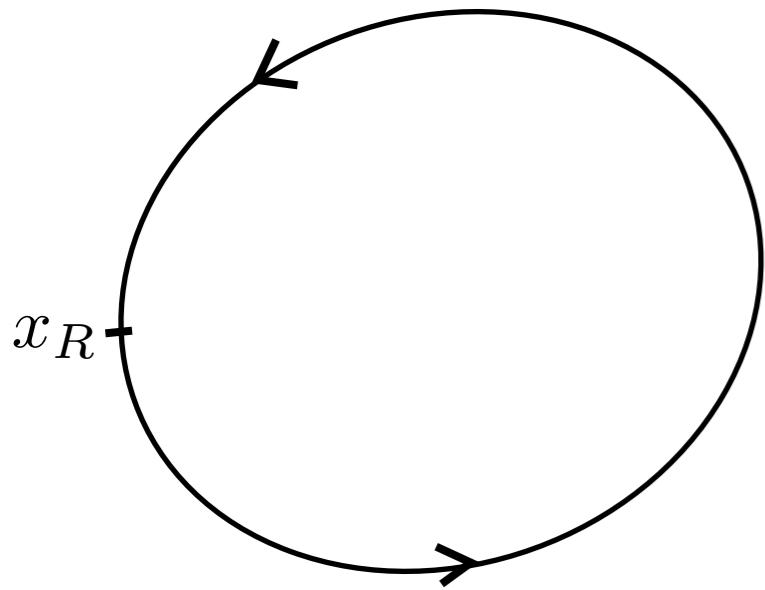


Path in loop space

# Path independency

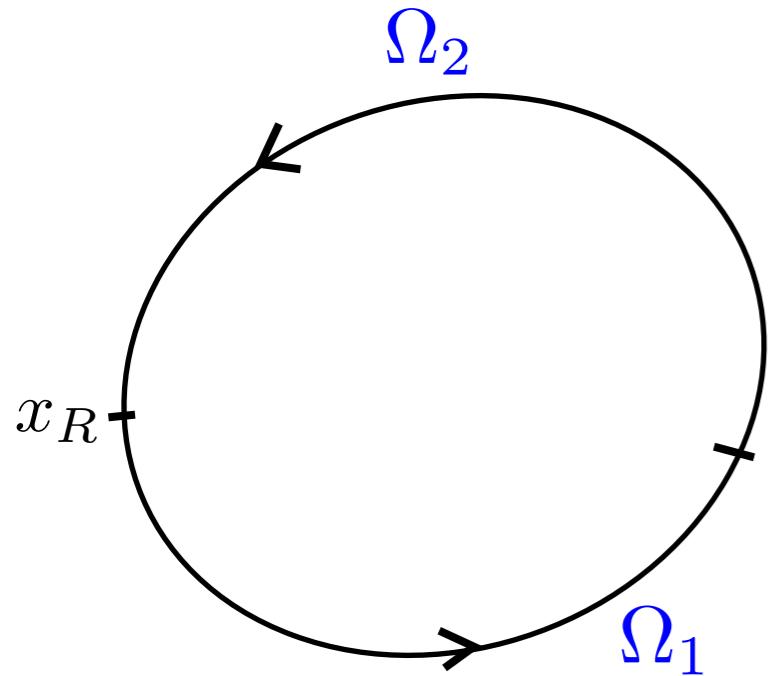
Closed volume (path)

## Path independency



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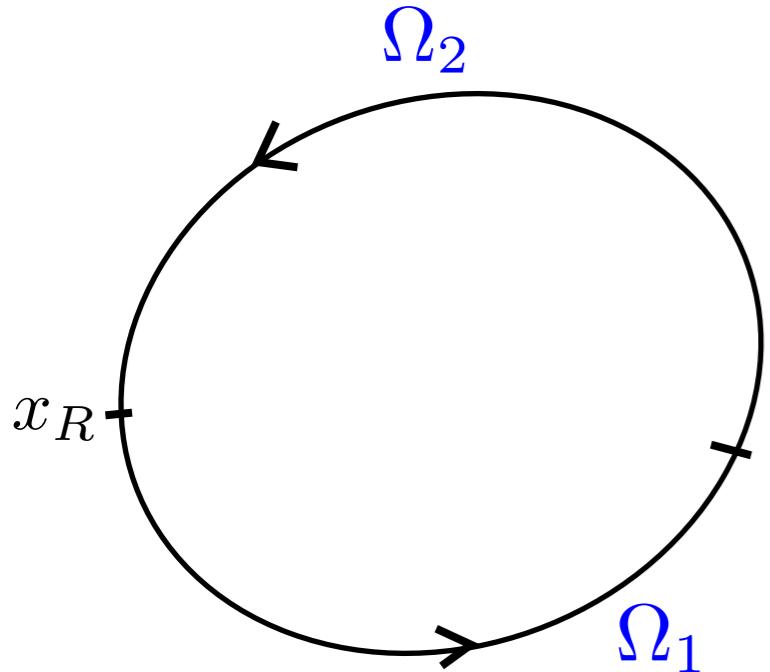
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$$\Omega_c = \Omega_1 + \Omega_2$$

Closed volume (path)

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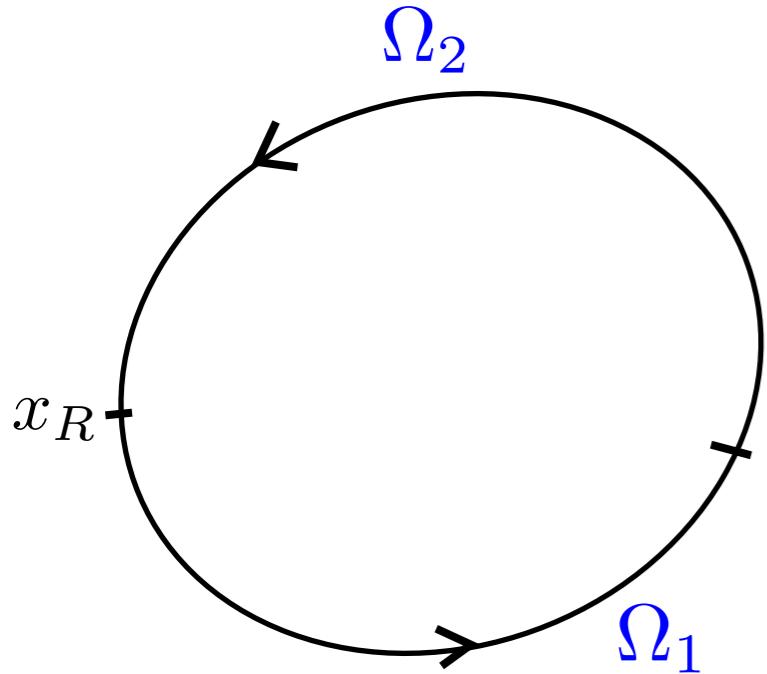


$$\Omega_c = \Omega_1 + \Omega_2$$

$$1 = P_3 e^{\oint_{\Omega_c} d\zeta d\tau V \mathcal{J} V^{-1}} = P_3 e^{\oint_{\Omega_2} d\zeta d\tau V \mathcal{J} V^{-1}} P_3 e^{\oint_{\Omega_1} d\zeta d\tau V \mathcal{J} V^{-1}}$$

Closed volume (path)

## Path independency



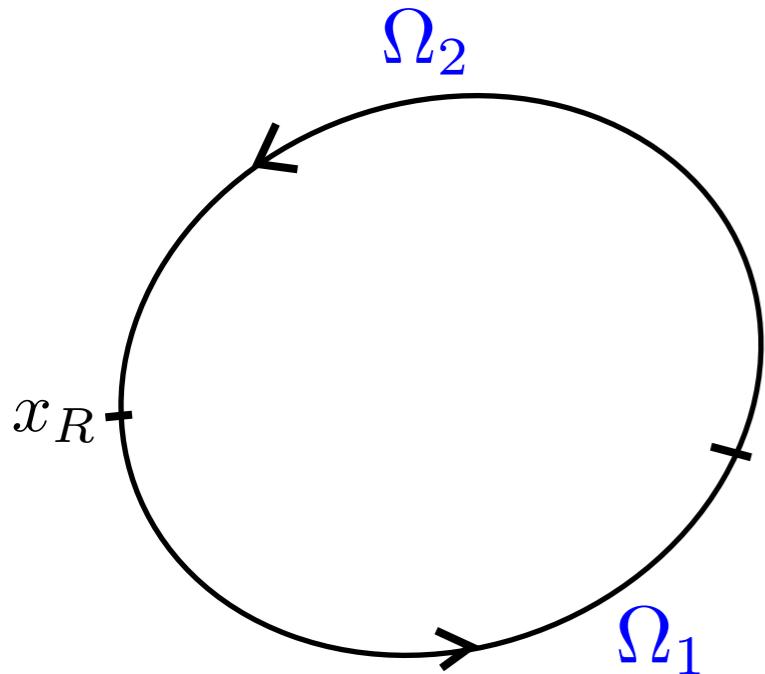
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$\Omega_1$  and  $\Omega_2^{-1}$  are volumes (paths) with the same end points

Closed volume (path)

## Path independency

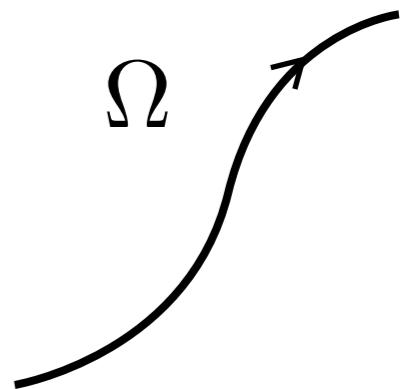


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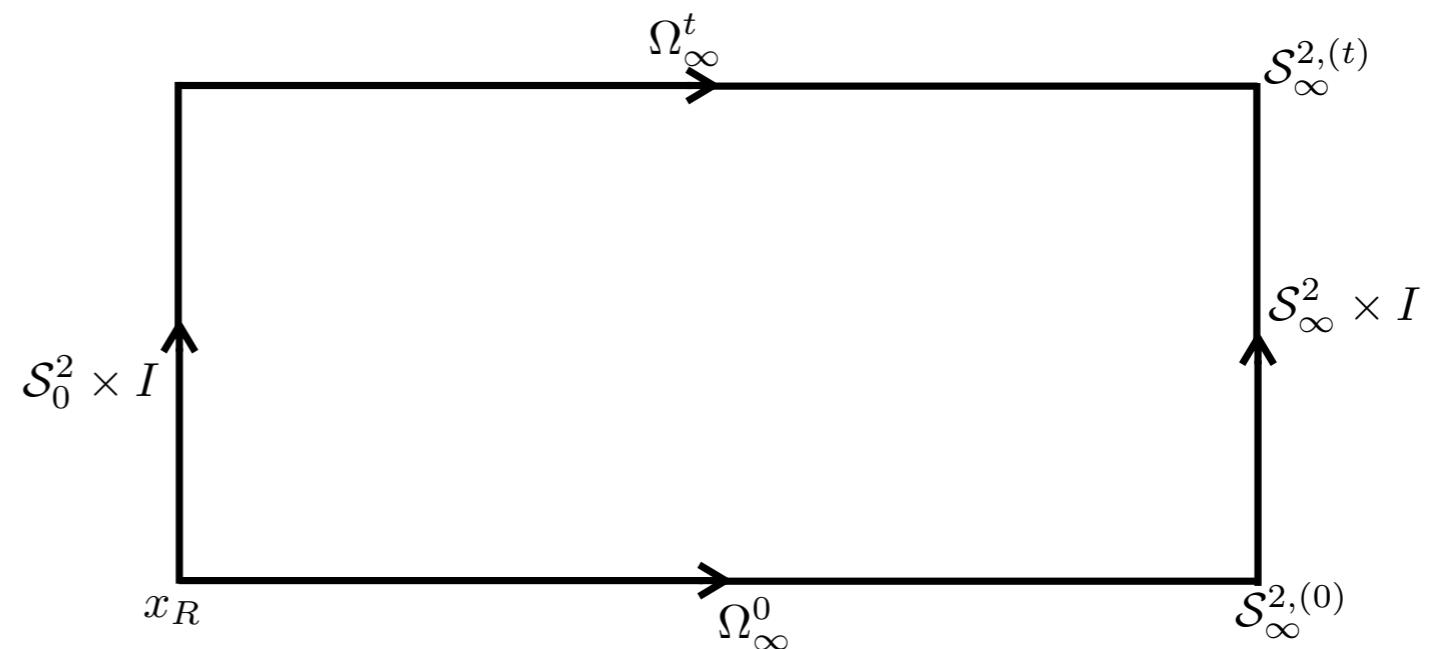
So,  $P_3 e^{\oint_{\Omega} d\zeta d\tau V \mathcal{J} V^{-1}}$  is path independent



# Conserved charges

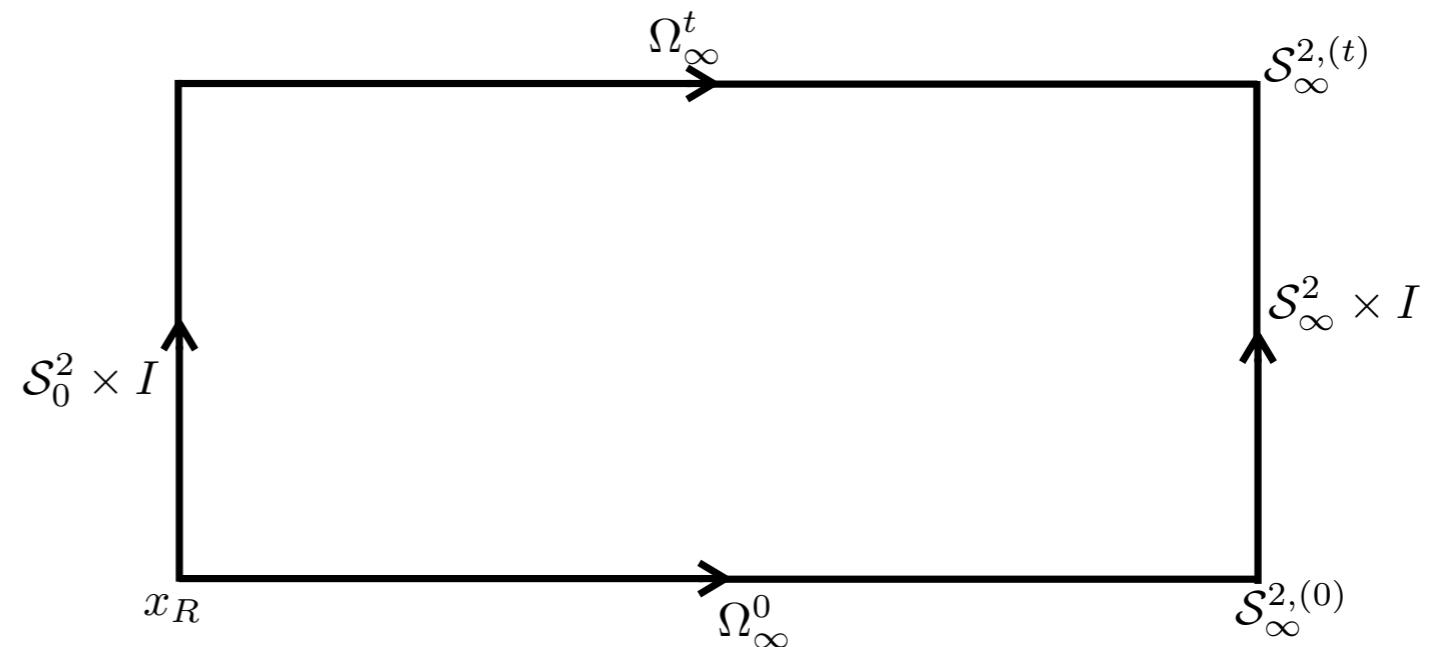
# Conserved charges

Two paths in loop space  
or volumes in space-time



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Two paths in loop space  
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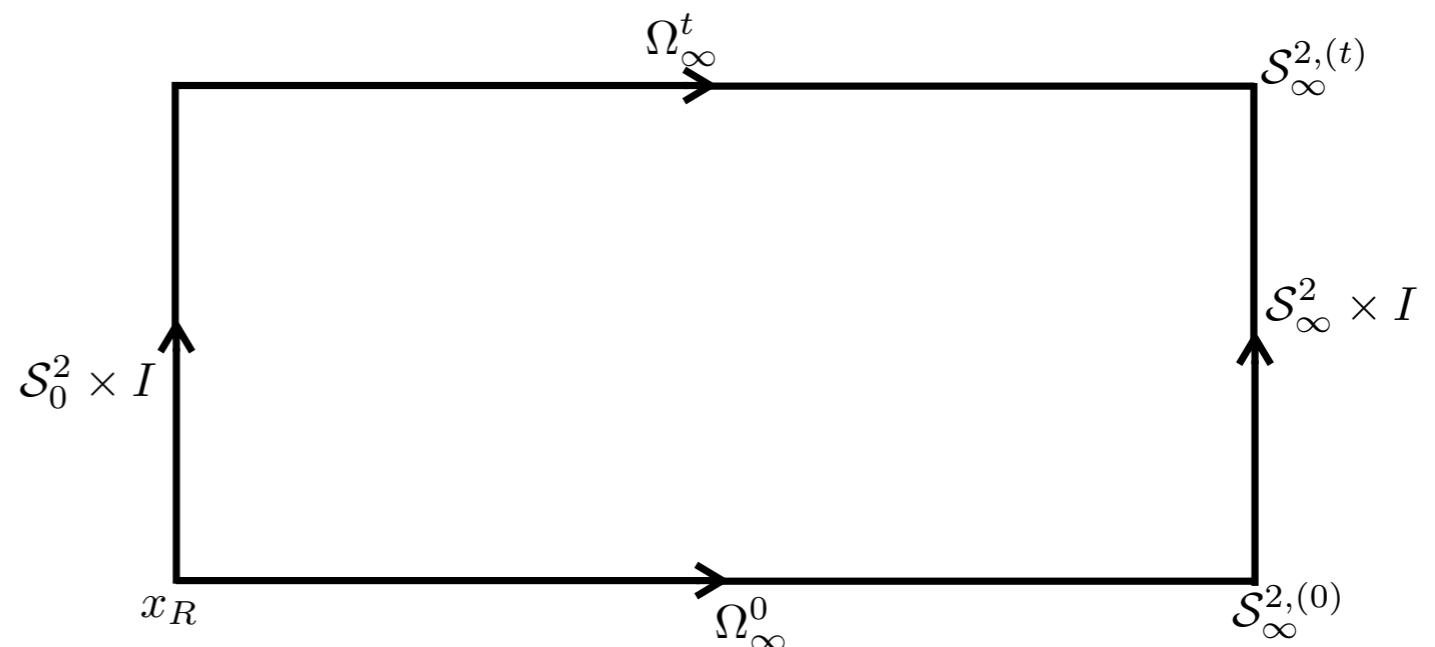
Path independency  $\rightarrow$

$$V(\mathcal{S}_\infty^{(2)} \times I) V(\Omega_\infty^{(0)}) = V(\Omega_\infty^{(t)}) V(\mathcal{S}_0^2 \times I)$$

$$\left[ V(*) \equiv P_3 e^{\int_* d\zeta d\tau V \mathcal{J} V^{-1}} \right]$$

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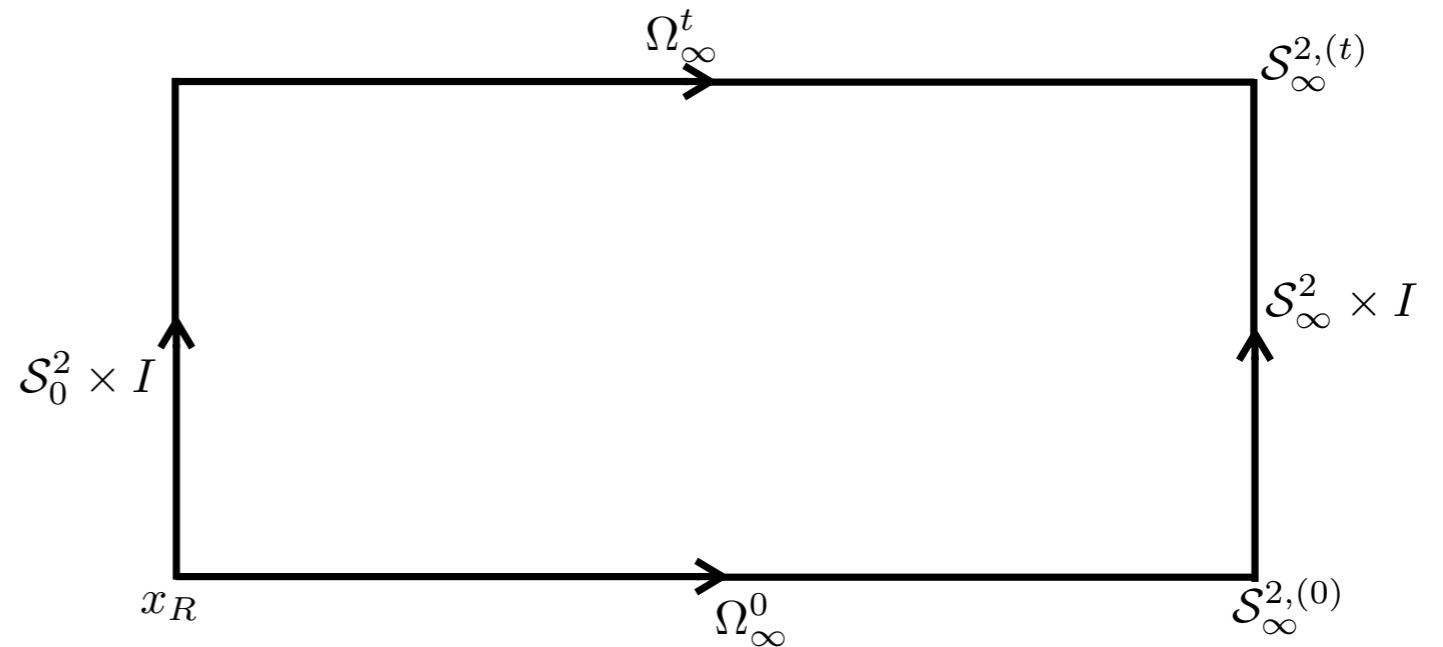
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The boundary conditions

$$J_\mu \sim \frac{1}{R^{2+\delta}} \quad F_{\mu\nu} \sim \frac{1}{R^{\frac{3}{2}+\delta'}}$$

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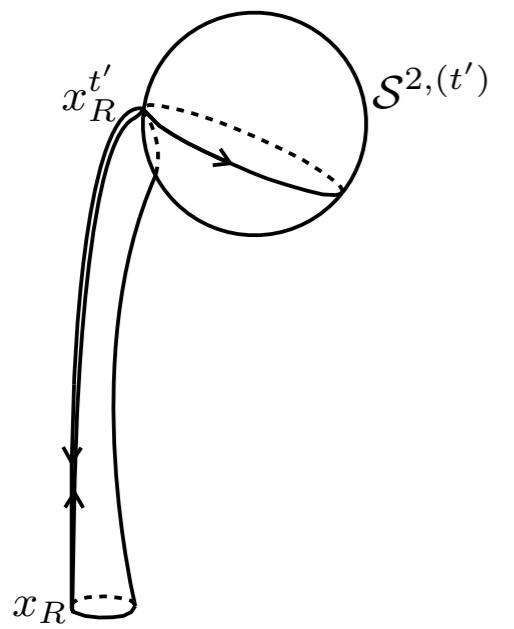
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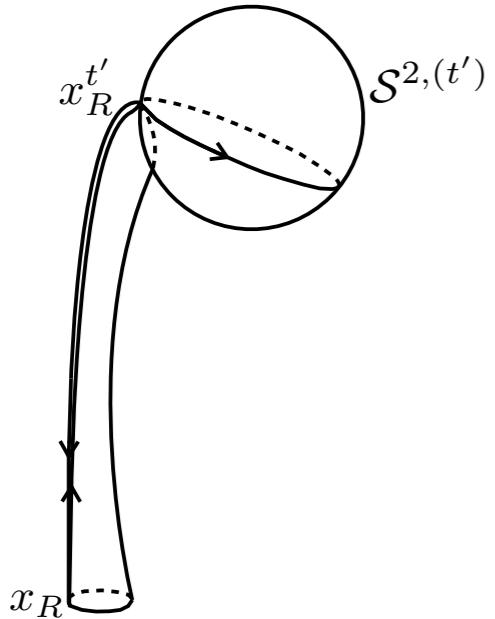
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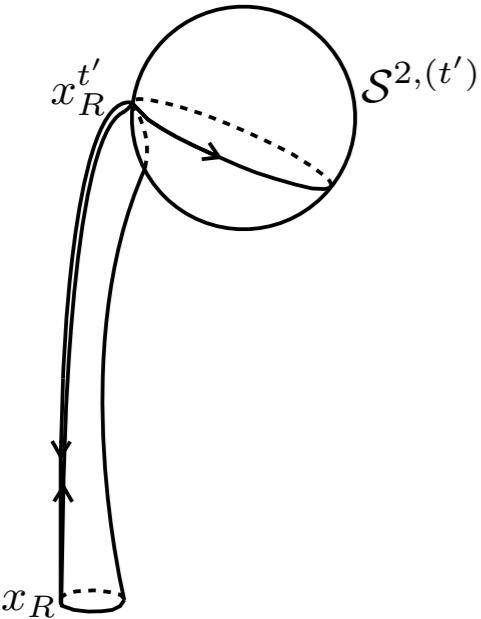
imply

$$V(\mathcal{S}_\infty^{(2)} \times I) = V(\mathcal{S}_0^2 \times I)$$





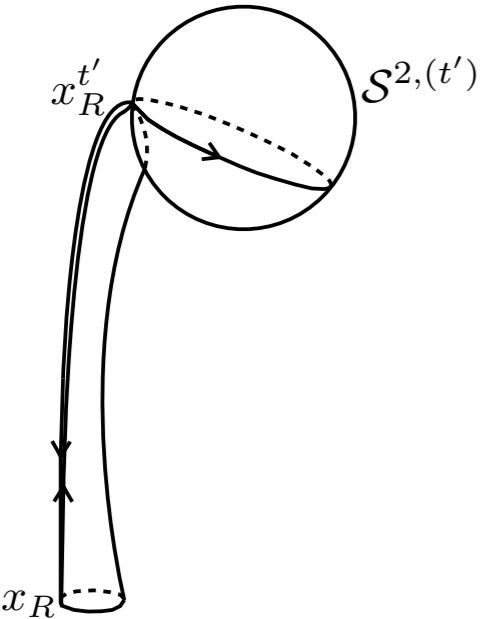
$$V_{x_R}(\Omega_\infty^{(t)}) = W^{-1}(x_R^{(t)}, x_R) V_{x_R^{(t)}}(\Omega_\infty^{(t)}) W(x_R^{(t)}, x_R)$$



$$V_{x_R}(\Omega_\infty^{(t)}) = W^{-1}(x_R^{(t)}, x_R) V_{x_R^t}(\Omega_\infty^{(t)}) W(x_R^{(t)}, x_R)$$

## Iso-spectral evolution

$$V_{x_R^t}(\Omega_\infty^{(t)}) = U(t) \cdot V(\Omega_\infty^{(0)}) \cdot U^{-1}(t)$$



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## Iso-spectral evolution

$$V_{x_R^{(t)}}(\Omega_\infty^{(t)}) = U(t) \cdot V(\Omega_\infty^{(0)}) \cdot U^{-1}(t)$$

with

$$U(t) = W(x_R^{(t)}, x_R) \cdot V(\mathcal{S}_0^2 \times I)$$

Conserved charges are eigenvalues of the operator

$$V_{x_R^{(t)}}(\Omega_\infty^{(t)}) = P_2 e^{ie \int_{\mathcal{S}_\infty^{2,(t)}} d\tau d\sigma (\alpha F_{\mu\nu}^W + \beta \tilde{F}_{\mu\nu}^W) \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}} = P_3 e^{\int_{\Omega_\infty^{(t)}} d\zeta d\tau V \mathcal{J} V^{-1}}$$

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2) Independent of reference point

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$P_3 e^{\int_{\Omega_\infty^{(t)}} d\zeta d\tau V \mathcal{J} V^{-1}}$  is path independent

# Example: BPS dyon

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Gauge theory with symmetry broken from  $G$  to  $H$   
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## Example: BPS dyon

Gauge theory with symmetry broken from  $G$  to  $H$   
by a Higgs  $\phi$  in the adjoint representation

$$\text{BPS eqs} \rightarrow E_i = -F^{0i} = \sin \theta D_i \phi \quad B_i = -\frac{1}{2} \epsilon^{ijk} F_{jk} = \cos \theta D_i \phi$$

$$D^i \phi \rightarrow \frac{\hat{r}^i}{4\pi r^2} G(\hat{r}) \quad D_\mu G(\hat{r}) = 0$$

For  $r \rightarrow \infty$

$$A_\mu = \frac{i}{e} \partial_\mu W W^{-1} \rightarrow G(\hat{r}) = W G(x_R) W^{-1}$$

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Conserved charges are eigenvalues of  $G(x_R)$

## The self-dual sector

$$F_{\mu\nu} = \kappa \tilde{F}_{\mu\nu} \quad \tilde{F}_{\mu\nu} = \frac{1}{2}\varepsilon_{\mu\nu\rho\sigma}F^{\rho\sigma} \quad \kappa = \pm 1$$

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Integral form is given by

$$P_1 e^{-ie \oint_{\partial\Sigma} d\sigma A_\mu \frac{dx^\mu}{d\sigma}} = P_2 e^{ie \int_\Sigma d\sigma d\tau W^{-1} [\alpha F_{\mu\nu} + \kappa (1-\alpha) \tilde{F}_{\mu\nu}] W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}}$$

with  $\Sigma$  being any two-dimensional surface on the space-time,  
and  $\alpha$  being an arbitrary parameter.

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$$P_1 e^{- \oint_{\partial\Sigma} d\sigma C_\mu \frac{dx^\mu}{d\sigma}} W_R = W_R P_2 e^{\int_\Sigma d\sigma d\tau W^{-1} H_{\mu\nu} W \frac{dx^\mu}{d\sigma} \frac{dx^\nu}{d\tau}}$$

**with**  $H_{\mu\nu} = \partial_\mu C_\nu - \partial_\nu C_\mu + [C_\mu, C_\nu]$

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$$H_{\mu\nu} = ie F_{\mu\nu} \xrightarrow{\hspace{10cm}} F_{\mu\nu} = \kappa \tilde{F}_{\mu\nu}$$

# Conserved Charges

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The charges are the eigenvalues of the operator

$$\begin{aligned} V_{x_R^{(t)}} \left( \mathcal{D}_{\infty}^{(t)} \right) &= P_2 e^{ie \int_{\mathcal{D}_{\infty}^{(t)}} d\sigma d\tau W^{-1} [\alpha F_{\mu\nu} + \kappa (1-\alpha) \tilde{F}_{\mu\nu}] W \frac{dx^{\mu}}{d\sigma} \frac{dx^{\nu}}{d\tau}} \\ &= P_1 e^{-ie \oint_{S_{\infty}^{1,(t)}} d\sigma A_{\mu} \frac{dx^{\mu}}{d\sigma}} \end{aligned}$$

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if the boundary conditions are satisfied

$$F_{\rho\sigma} = \kappa \tilde{F}_{\rho\sigma} \sim \frac{1}{r^{2+\delta}} T(\hat{r}) \quad \text{for} \quad r \rightarrow \infty$$

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$$A_\mu = -\frac{2}{e} \sigma_{\mu\nu} \frac{x^\nu - a^\nu}{(x^\rho - a^\rho)^2 + \lambda^2}$$

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One gets

$$P_1 e^{-ie \oint_{S_\infty^{1,(t)}} d\sigma A_\mu \frac{dx^\mu}{d\sigma}} = e^{-i 2 \int_0^{2\pi} d\theta \sigma_{\mu\nu}} = e^{-i 4\pi \sigma_{\mu\nu}}$$

# Are integrable theories gauge theories?

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Look for theories in  $d + 1$  dimensions where the equations of motion take the integral form

$$P_{d-1} e^{\int_{\partial\Omega} \mathcal{F}} = P_d e^{\int_{\Omega} \mathcal{J}}$$

with  $\Omega$  a  $d$ -dimensional hyper volume

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Examples:

- 1) Integrable theories in  $1 + 1$  dimensions (soliton theories)
- 2) Chern-Simons theories in  $2 + 1$  dimensions
- 3) Yang-Mills in  $2 + 1$  and  $3 + 1$  dimensions

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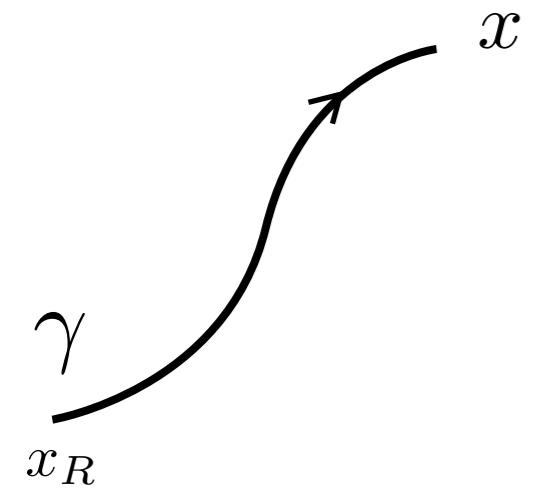
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L. A. Ferreira and G. Luchini,

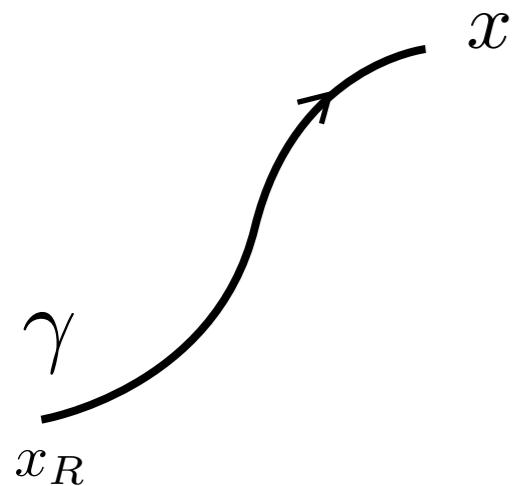
An integral formulation of Yang-Mills on loop space, [arXiv:1109.2120 [hep-th]]  
Gauge and Integrable Theories in Loop Spaces, [arXiv:1109.2606 [hep-th]],  
Nuclear Physics B 858 [PM] (2012) 336–365

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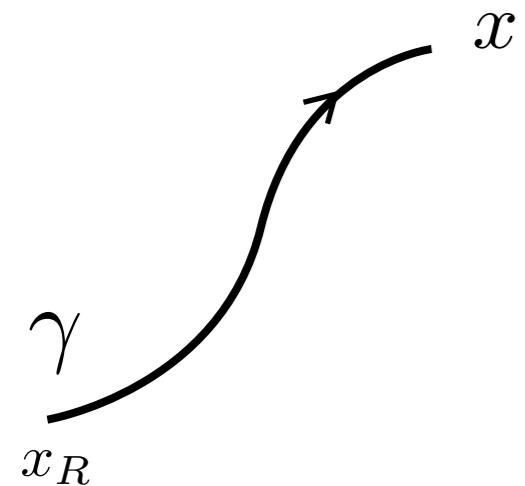


## In $(1+1)$ -dimensions the volume is one-dimensional



Let  $g(x)$  and  $A_\mu$  be the fields of the theory  
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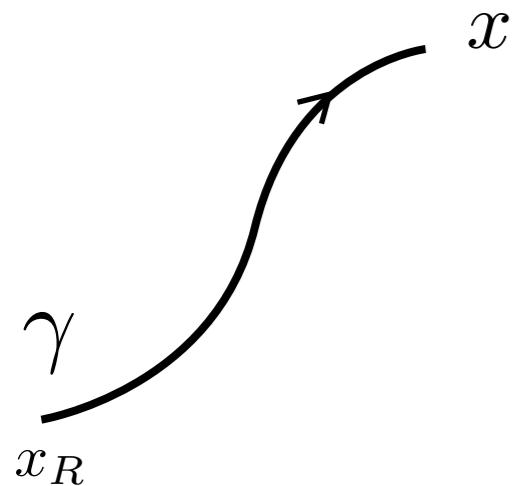
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The fact that it must be true for any curve joining  $x_R$  to  $x$ , it implies  $A_\mu$  is flat

$$\partial_\mu A_\nu - \partial_\nu A_\mu + [A_\mu, A_\nu] = 0 \quad A_\mu = -\partial_\mu g g^{-1}$$

Thank You