

# Holographic superconductors and superfluids: effect of backreaction

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Introduction  
Does it work?  
The model

Holographic superconductors in 3+1 dimensions  
Holographic superfluids in 2+1 dimensions  
Summary



Distance São Carlos – Bremen:  
≈ 9980 km (direct line)

Population of Bremen:  
≈ 550.000  
(10<sup>th</sup> biggest city in Germany)



# Collaborations and References

Work done in collaboration with:

**Yves Brihaye** - *Université de Mons, Belgium*

References:

- Y. Brihaye and B. Hartmann, Phys. Rev. D81 (2010) 126008*
- Y. Brihaye and B. Hartmann, JHEP 1009 (2010) 002*
- Y. Brihaye and B. Hartmann, Phys. Rev. D83 (2011) 126008*

# Outline

- 1 Introduction
- 2 Does it work?
- 3 The model
- 4 Holographic superconductors in 3+1 dimensions
- 5 Holographic superfluids in 2+1 dimensions
- 6 Summary

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# The holographic principle

## The AdS/CFT correspondence (Maldacena, 1997)

A theory of **gravity** in  $d$ -dimensional asymptotically Anti-de Sitter ( $\text{AdS}_d$ ) space-time is dual to a **conformal field theory** (CFT) living on the  $(d - 1)$ -dimensional boundary of  $\text{AdS}_d$ .

- important result of String Theory
- **weak coupling  $\Leftrightarrow$  strong coupling duality**
- CFT is **scale-invariant** Quantum field theory (QFT)
- $\text{AdS}_d$ : Vacuum solution of  $d$ -dimensional Einstein's equations with **negative cosmological constant**

# Applying the AdS/CFT correspondence

use classical gravity theory (weakly-coupled) to study strongly coupled Quantum Field theories

- gravity  $\Leftrightarrow$  condensed matter (Quantum phase transitions):

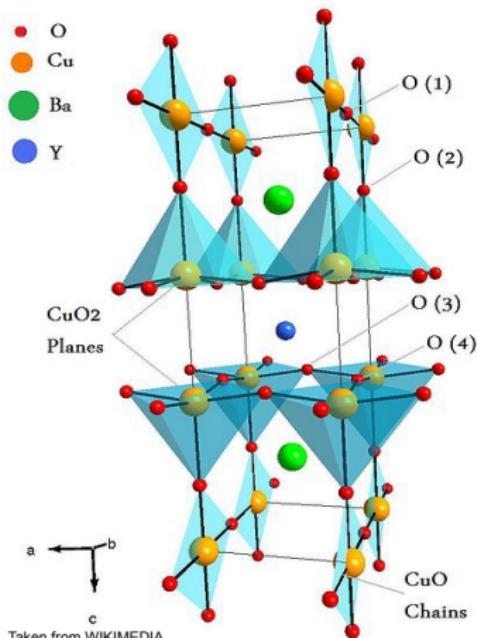
## **holographic superconductors/superfluids**

- Superconductor/insulator phase transitions in thin metallic films
- high temperature superconductors
- ...
- gravity  $\Leftrightarrow$  Quantumchromodynamics (QCD) (e.g. quark-gluon plasma ...)

# Quantum Phase Transitions

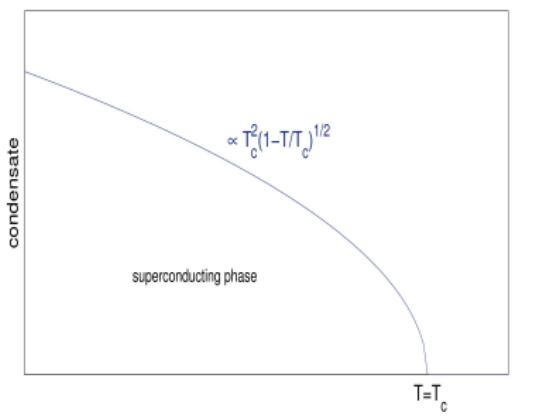
- phase transition (PT) at  $T = 0$
- not **thermal**, but **quantum** fluctuations  
(uncertainty principle)
- PT at critical parameter, e.g. at chemical potential  $\mu = \mu_c$
- Quantum critical region at  $\mu \approx \mu_c$  and  $T > 0$
- **believed to appear in high- $T$  superconductors**
- (believed to be) described by **strongly coupled** theories  
→ “standard” (mainly weakly coupled) theories  
(BCS, Ginzburg-Landau) do **not** work well

# Example of a high temperature superconductor



- Yttrium(Y)-barium(Ba)-copper(Cu)-oxide(O)
- highest possible  $T_c = 92\text{K}$  (boiling point of liquid nitrogen: 77K)
- **superconductivity associated to CuO<sub>2</sub>-planes**

# Holographic phase transitions



## Basic ingredients

- notion of **temperature**  
⇒ black hole
  - notion of **chemical potential**  
⇒ **charged** black hole
  - notion of **condensate**  
⇒ non-trivial field outside  
black hole horizon
- black hole for  $T > T_c$   
“hairy” black hole for  $T < T_c$

# Properties of black holes

- **temperature**  $T = \frac{\kappa}{2\pi}$  with surface gravity

$$\kappa^2 = -\frac{1}{2}(D_\mu \chi_\nu)(D^\mu \chi^\nu) \Big|_{r_h}$$

where  $\chi$  Killing

- **free energy**  $\Omega$  ( $\rightarrow$  canonical ensemble)

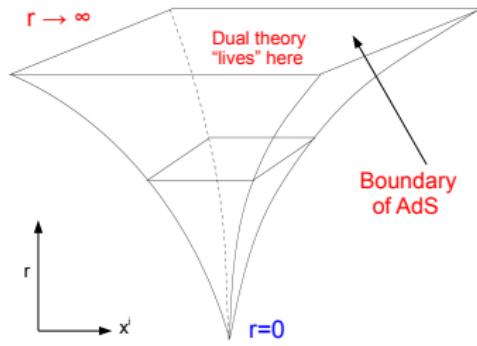
$$\Omega = TS_{\text{os}}$$

$S_{\text{os}}$ : action evaluated on-shell

# Planar Anti-de Sitter (AdS) space-time

- Metric in (3+1) dimensions

$$ds^2 = -\frac{r^2}{\ell^2} dt^2 + \frac{\ell^2}{r^2} dr^2 + \frac{r^2}{\ell^2} (dx^2 + dy^2)$$



Taken from arxiv: 0808.1115

- cosmological constant

$$\Lambda = -3/\ell^2$$

- Ricci scalar

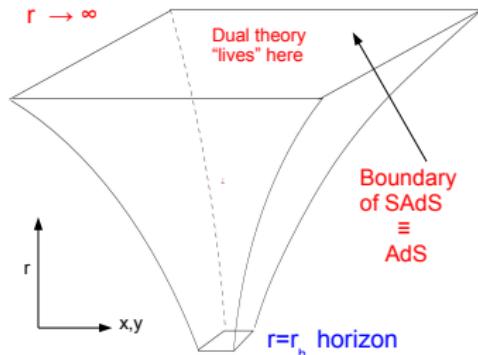
$$R = -12/\ell^2$$

$\Rightarrow \ell$  AdS radius

# Planar Schwarzschild-Anti-de Sitter (SAdS)

- Metric in (3+1) dimensions

$$ds^2 = - \left( \frac{r^2}{\ell^2} - \frac{r_h^3}{r\ell^2} \right) dt^2 + \left( \frac{r^2}{\ell^2} - \frac{r_h^3}{r\ell^2} \right)^{-1} dr^2 + \frac{r^2}{\ell^2} (dx^2 + dy^2)$$



Taken from arxiv: 0808.1115

- Black hole with planar horizon at  $r = r_h$
- temperature**  $T = 3r_h/(4\pi\ell^2)$
- asymptotically  $\text{AdS}_{3+1}$

# Planar Reissner-Nordström-Anti-de Sitter (RNAdS)

- Metric in (3+1) dimensions

$$ds^2 = -f(r)dt^2 + \frac{1}{f(r)}dr^2 + \frac{r^2}{\ell^2}(dx^2 + dy^2)$$

with

$$f(r) = \frac{r^2}{\ell^2} - \frac{1}{r} \left( \frac{\textcolor{blue}{r}_+^3}{\ell^2} + \frac{q^2}{4\textcolor{blue}{r}_+} \right) + \frac{q^2}{4r^2}$$

- U(1) gauge field

$$A_M dx^M = A_t(r)dt = \frac{q}{\textcolor{blue}{r}_+} - \frac{q}{r} = \textcolor{red}{\mu} \left( 1 - \frac{\textcolor{blue}{r}_+}{r} \right)$$

- planar (outer) horizon at  $r = \textcolor{blue}{r}_+$ , asymptotically AdS<sub>3+1</sub>
- $q \propto$  charge density,  $\mu$  chemical potential
- temperature**  $T = 3\textcolor{blue}{r}_+/(4\pi\ell^2) - \textcolor{red}{\mu}^2/(4\textcolor{blue}{r}_+)$

# Anti-de Sitter soliton

- **double Wick rotation** ( $t \rightarrow i\eta$ ,  $y \rightarrow it$ ) of SAdS
- metric in (3+1) dimensions

$$ds^2 = -\frac{r^2}{\ell^2} dt^2 + \left( \frac{r^2}{\ell^2} - \frac{r_0^3}{r\ell^2} \right)^{-1} dr^2 + \left( \frac{r^2}{\ell^2} - \frac{r_0^3}{r\ell^2} \right) d\eta^2 + \frac{r^2}{\ell^2} dx^2$$

- to avoid conical singularity  $\rightarrow \eta$  **periodic** with period

$$\tau_\eta = \frac{4\pi\ell^2}{3r_0} \quad \text{where } r_0 > 0$$

- space-time unchanged for  $A_M dx^M = \mu dt$
- can be considered at **any temperature**  $T$
- **unstable** to decay to SAdS for large  $T$   
(Surya, Schleich & Witt, 2001)

# Formation of scalar hair on AdS black hole

- Action ( $M, N = 0, \dots, d - 1$ )

$$S = \int d^d x \sqrt{-g} \left[ \frac{R - 2\Lambda}{16\pi G} - (\partial_M \psi)^* \partial^M \psi - m^2 \psi^* \psi \right]$$

$\psi$ : complex scalar field

$m$ : mass of scalar field

$G$ : Newton's constant

$\Lambda$ : (negative) cosmological constant

- **Breitenlohner-Freedman (BF) bound:**  $\text{AdS}_d$  stable against scalar hair formation for

$$m^2 \geq m_{\text{BF}}^2 = -\frac{(d-1)^2}{4\ell^2}$$

(Breitenlohner & Freedman, 1982)

# Formation of scalar hair on AdS black hole

- Action ( $M, N = 0, \dots, d - 1$ )

$$S = \int d^d x \sqrt{-g} \left[ \frac{R - 2\Lambda}{16\pi G} - (D_M \psi)^* D^M \psi - m^2 \psi^* \psi - \frac{1}{4} F_{MN} F^{MN} \right]$$

$F_{MN} = \partial_M A_N - \partial_N A_M$ : field strength of U(1) gauge field  $A_M$

$D_M \psi = \partial_M \psi - ie A_M \psi$ : covariant derivative

e: gauge coupling

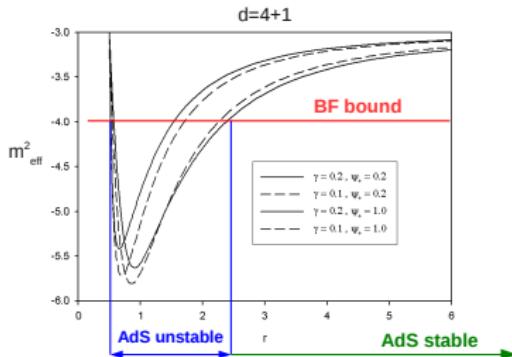
- effective mass of scalar field for  $A_i = 0, i = 1, \dots, d - 1$

$$m_{\text{eff}}^2 = m^2 - e^2 |g^{tt}| A_t^2$$

- $e^2 |g^{tt}|$  sufficiently large  $\Rightarrow$  **unstable**

# Formation of scalar hair on AdS black hole

- Example (Y. Brihaye & B.Hartmann, Phys.Rev. D81 (2010) 126008)



- Pioneering example: RNAdS black hole close to  $T = 0$  unstable to form **scalar hair** close to horizon (Gubser, 2008)

# Formation of scalar hair on AdS soliton

- ground state energy of scalar field in AdS soliton background  
**finite and positive**  $\Rightarrow$  **energy gap** (Witten, 1998)
- coupling to U(1) gauge field:

$$A_M dx^M = \phi(r) dt \quad \text{with} \quad \phi(r \rightarrow \infty) \rightarrow \mu$$

**decreases** ground state energy

- for  $T$  small and  $\mu > \mu_c$  soliton **unstable to form scalar hair**  
(Nishioka, Ryu & Takayanagi, 2010; Horowitz & Way, 2010;  
Brihaye & B. Hartmann, 2011)

# AdS/CFT “dictionary” in (3+1) dimensions

(Hartnoll, Herzog & Horowitz, 2008)

- formation of scalar hair on (3+1)-dimensional AdS black hole/soliton dual to onset of superconductivity in (2+1) dimensions
- U(1) gauge field dual to global U(1) field with

$$A_t(r \gg 1) = \mu - \frac{q}{r} + \dots$$

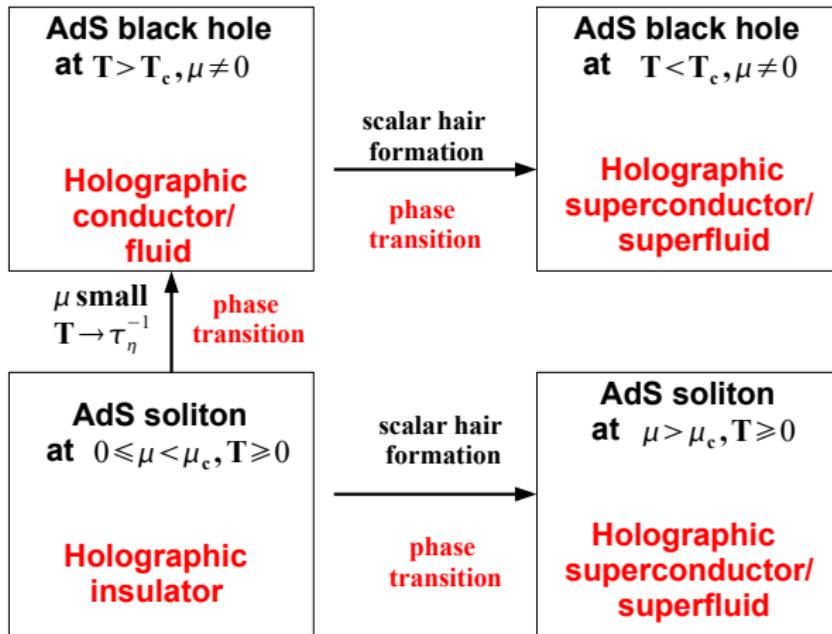
$\mu$ : dual to **chemical potential** ,  $q$ : dual to charge density

- scalar field with  $m^2 > m_{\text{BF}}^2 = -9/(4\ell^2)$ :

$$\psi(r \gg 1) = \frac{\psi_-}{r} + \frac{\psi_+}{r^2} + \dots \quad \text{for } m^2 = -2/\ell^2$$

$\psi_{\pm}$ : dual to expectation value of operator  $\mathcal{O}_{\pm}$   
 $\Rightarrow$  **value of condensate**

# Holographic phase transitions

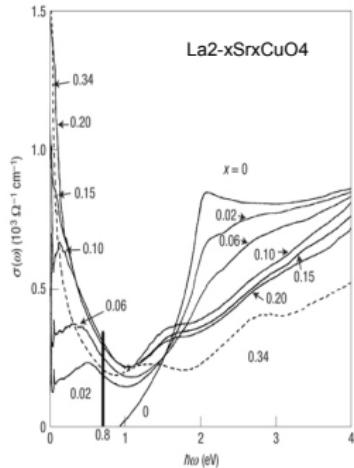


# Outline

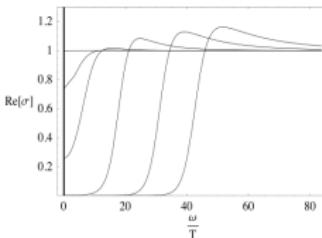
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# First results on holographic superconductors

- optical conductivity: conductivity  $\sigma$  as function of frequency  $\omega$



Experimental  
Uchida et. al (1991)



Theoretical  
Hartnoll, Herzog, Horowitz (2008)

# First results on holographic superconductors

- **London equation**  $\vec{E} \propto \frac{\partial \vec{J}}{\partial t}$  fulfilled
- boundary magnetic fields not dynamical:  
**Meissner effect** difficult to model
- Correlation length  $\xi T_c \approx 0.1(1 - T/T_c)^{-1/2}$   
(Horowitz, and Roberts, 2008)
- Holographic superconductors are type II  
... just like most high-T superconductors...  
⇒ formation of **vortices** with quantized magnetic flux  
(Montull, Pomarol, and Silva, 2009)

# First results on holographic superconductors

- **frequency gap**  $\omega_g \doteq$  energy required to break Cooper pair
  - from BCS theory:  
 $\omega_g/T_c \approx 3.5$
  - Holographic superconductors:  
 $\omega_g/T_c \approx 8$   
(Hartnoll, Herzog, Horowitz, 2008)
  - compare to **experimental**  $Bi_2Sr_2CaCu_2O_{8+x}$  values:  
 $\omega_g/T_c = 7.9 \pm 0.5$   
(Gomes et al., 2007)

# One problem though...

- **Mermin-Wagner theorem:** spontaneous symmetry breaking forbidden in (2+1) dimensions at finite temperature, but holographic superconductors have been constructed
- BUT: Einstein gravity (...used mostly..) corresponds to large  $N$  limit on QFT side

**Q: Can higher curvature corrections (e.g. Gauss-Bonnet terms) suppress condensation?**

# The model

**Gauss-Bonnet gravity** in d-dimensional Anti-de Sitter ( $\text{AdS}_d$ )

$$S = \frac{1}{16\pi G} \int d^d x \sqrt{-g} \left( R - 2\Lambda + \frac{\alpha}{2} \mathcal{L}_{\text{GB}} + 16\pi G \mathcal{L}_{\text{m}} \right) + S_{\text{ct}}$$

Gauss-Bonnet Lagrangian

$$\mathcal{L}_{\text{GB}} = (R^{MNKL}R_{MNKL} - 4R^{MN}R_{MN} + R^2)$$

$S_{\text{ct}}$ : boundary **counterterm** - necessary to make action finite

$G$ : Newton's constant

$\Lambda = -(d-1)(d-2)/(2\ell^2)$ : cosmological constant

$\alpha$ : Gauss–Bonnet coupling

# The model

Lagrangian of **charged complex scalar field**:

$$\mathcal{L}_m = -\frac{1}{4}F_{MN}F^{MN} - (D_M\psi)^* D^M\psi - m^2\psi^*\psi , \quad M, N = 0, 1, 2, 3, d-1$$

U(1) field strength tensor

$$F_{MN} = \partial_M A_N - \partial_N A_M$$

covariant derivative

$$D_M\psi = \partial_M\psi - ieA_M\psi$$

$e$ : gauge coupling

$m$ : mass of scalar field

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# The Ansatz for AdS black holes

- Metric

$$ds^2 = -f(r)a^2(r)dt^2 + \frac{1}{f(r)}dr^2 + r^2(dx^2 + dy^2 + dz^2)$$

with  $f(r_h) = 0$  at **horizon**  $r = r_h$

- Electric field only

$$A_M dx^M = \phi(r)dt$$

- Gauge freedom: scalar field chosen to be **real**

$$\psi = \psi(r)$$

# Parameters

- Equations invariant under **rescalings**

$$\psi \rightarrow \psi/e , \quad \phi \rightarrow \phi/e , \quad G \rightarrow e^2 G$$

$$r \rightarrow \ell r , \quad (t, x, y, z) \rightarrow (t, x, y, z)/\ell , \quad f \rightarrow \ell^2 f , \quad \phi \rightarrow \ell \phi$$

$\Rightarrow e = \ell \equiv 1$  **without loss of generality**

- choose  $m^2 = -\frac{(d-2)}{\ell^2} > m_{\text{BF}}^2 = -\frac{(d-1)^2}{4\ell^2}$  for  $d = 4, 5$
- define  $\gamma = 8\pi G/e^2$

# Probe limit vs. Backreaction

- $G = 0$  ( $e = \infty$ ) “**probe limit**”: fixed space-time background  
⇒ **coupled scalar & gauge field equations**
- $G \neq 0$  ( $e < \infty$ ): **backreaction** of matter fields on space-time  
⇒ **coupled Einstein, scalar & gauge field equations**
- results in systems of **coupled, nonlinear ordinary or partial differential equations** that have to be solved **numerically**

# Equations of motion

$$\begin{aligned}
 f' &= 2r \frac{-f + 2r^2}{r^2 - 2\alpha f} \\
 &\quad - \gamma \frac{r^3}{2fa^2} \left( \frac{2\phi^2\psi^2 + f(2m^2a^2\psi^2 + \phi'^2) + 2f^2a^2\psi'^2}{r^2 - 2\alpha f} \right) \\
 a' &= \gamma \frac{r^3(\phi^2\psi^2 + a^2f^2\psi'^2)}{af^2(r^2 - 2\alpha f)} \\
 \phi'' &= - \left( \frac{3}{r} - \frac{a'}{a} \right) \phi' + 2 \frac{\psi^2}{f} \phi \\
 \psi'' &= - \left( \frac{3}{r} + \frac{f'}{f} + \frac{a'}{a} \right) \psi' - \left( \frac{\phi^2}{f^2a^2} - \frac{m^2}{f} \right) \psi
 \end{aligned}$$

# Conditions at the horizon $r = r_h$

- Horizon  $r = r_h$

$$f(r_h) = 0$$

- Regularity of matter fields on horizon

$$\phi(r_h) = 0 \quad , \quad \psi'(r_h) = \left. \frac{m^2 \psi r^2}{4r/\ell^2 - \gamma r^3 (m^2 \psi^2 + \phi'^2/(2a^2))} \right|_{r=r_h}$$

# Conditions on the AdS boundary $r \gg 1$

- Electric potential

$$\phi(r \gg 1) = \mu - q/r^2$$

$\mu$ : chemical potential

$q$ : charge density

- Scalar field

$$\psi(r \gg 1) = \frac{\psi_-}{r^{\lambda_-}} + \frac{\psi_+}{r^{\lambda_+}}$$

with

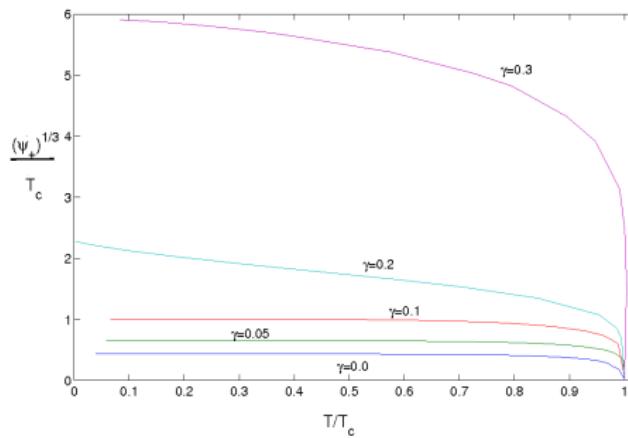
$$\lambda_{\pm} = 2 \pm \sqrt{4 - 3\tilde{\alpha}^2} , \quad \tilde{\alpha}^2 \equiv \frac{2\alpha}{1 - \sqrt{1 - 4\alpha}} , \quad \alpha \leq 1/4$$

$\psi_{\pm}$ : value of condensate in dual theory

# Holographic superconductors: backreaction

(Y. Brihaye & B.Hartmann, Phys.Rev. D81 (2010) 126008)

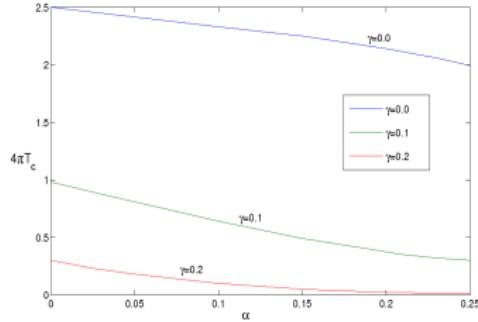
- Value of condensate increases with increasing  $\gamma$



# Holographic superconductors: critical temperature $T_c$

(Y. Brihaye & B.Hartmann, Phys.Rev. D81 (2010) 126008)

- temperature  $T_c$  of conductor/superconductor phase transition
  - for  $\alpha = 0$ :
$$T_c \approx 0.198 \cdot \exp(-10.6\gamma) q^{1/3}$$
  - for  $\alpha \neq 0$ :



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# The Ansatz for AdS black strings/AdS solitons

- $\alpha = 0$  in the following
- Metric of a (3+1)-dimensional **rotating AdS black string (BS)**

$$ds^2 = -b(\rho)dt^2 + \frac{1}{f(\rho)}d\rho^2 + \rho^2(g(\rho)dt - d\varphi)^2 + p(\rho)dz^2$$

with  $f(\rho_h) = b(\rho_h) = 0$  at horizon  $\rho = \rho_h$

- Metric of a (3+1)-dimensional **rotating AdS soliton (S)**

$$ds^2 = -p(\rho)dt^2 + \frac{1}{f(\rho)}d\rho^2 + b(\rho)(g(\rho)dt - d\eta)^2 + \rho^2dz^2$$

with  $f(\rho_0) = b(\rho_0) = 0$  and  $\eta$  **periodic** with period

$$\tau_\eta = \frac{4\pi}{\sqrt{b'(\rho_0)f'(\rho_0)}}$$

# The Ansatz

- U(1) gauge field for black strings

$$A_M dx^M = \phi(\rho) dt + A(\rho) d\varphi$$

- U(1) gauge field for AdS solitons

$$A_M dx^M = \phi(\rho) dt + A(\rho) d\eta$$

- Gauge freedom: scalar field chosen to be **real**

$$\psi = \psi(\rho)$$

# Conditions on the AdS boundary $\rho \gg 1$

- U(1) potential

$$\phi(\rho \gg 1) = \mu - q/\rho , \quad A(\rho \gg 1) = \sigma - \tilde{q}/\rho$$

$\mu$ : chemical potential

$\sigma$ : superfluid velocity

$q$ : electric charge density

$\tilde{q}$ : magnetic charge density

- Scalar field

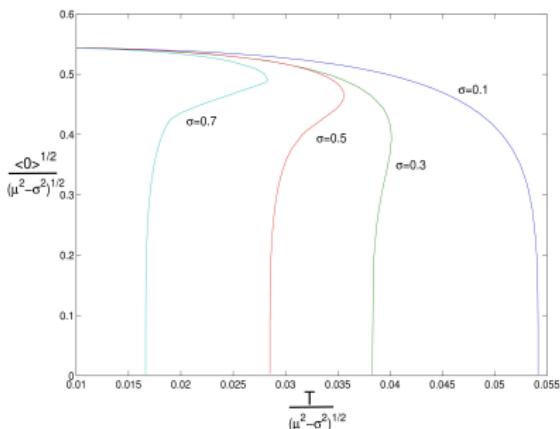
$$\psi(\rho \gg 1) = \frac{\psi_-}{\rho} + \frac{\psi_+}{\rho^2}$$

$\psi_{\pm}$ : value of condensate in dual theory

# Holographic superfluids: order of phase transition

(Y. Brihaye & B. Hartmann, JHEP 1009 (2010) 002 )

- Probe limit  $\gamma = 0$ , fluid/BS superfluid phase transition

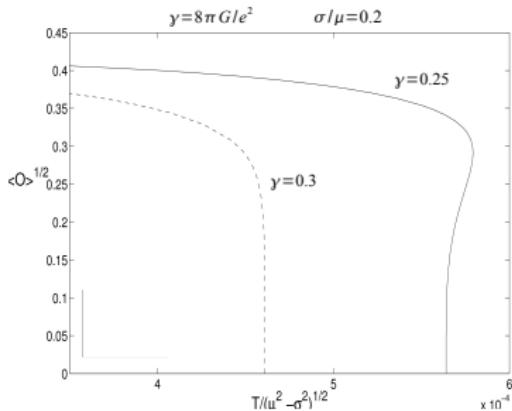


- 2nd order phase transition for  $\sigma$  small
- 1st order phase transition for  $\sigma$  large
- temperature  $T_c$  decreases with increasing  $\sigma$

# Holographic superfluids: order of phase transition

(Y. Brihaye & B. Hartmann, Phys. Rev. D83 (2011) 126008 )

- $\gamma \neq 0$ , fluid/BS superfluid phase transition

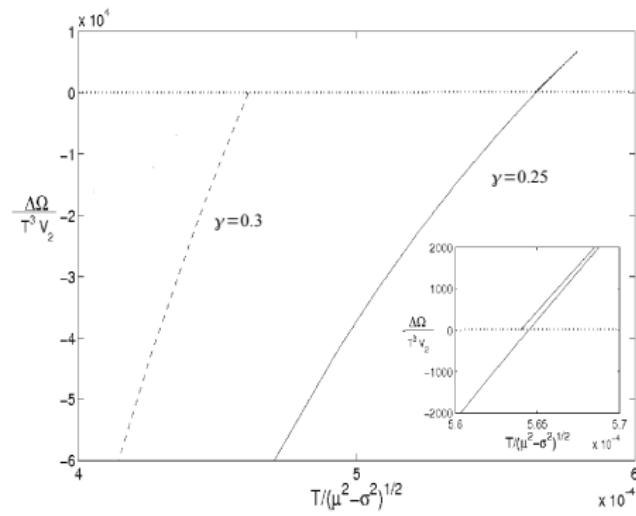


- 1st order for  $\sigma$  large and  $\gamma$  small
- 2nd order for  $\sigma \geq 0$  and  $\gamma$  large
- temperature  $T_c$  decreases with increasing  $\gamma$

# Holographic superfluids: free energy

(Y. Brihaye & B. Hartmann, Phys. Rev. D83 (2011) 126008)

- $\gamma \neq 0$ , fluid/BS superfluid phase transition



# Holographic superfluids: free energy

(Y. Brihaye & B. Hartmann, Phys. Rev. D83 (2011) 126008)

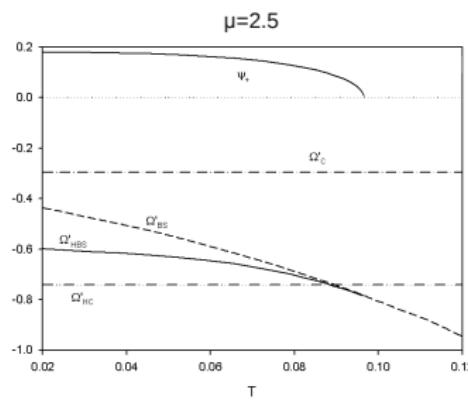
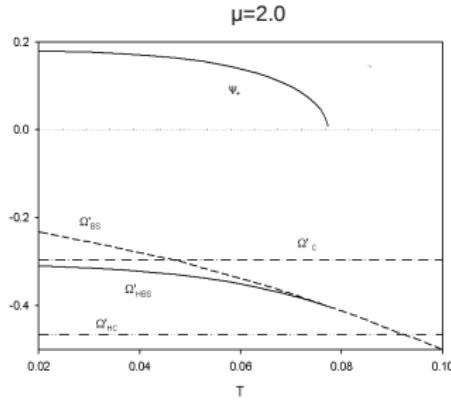
- **Free energy** of different phases with

BS: black string → fluid

HBS: hairy black string → superfluid

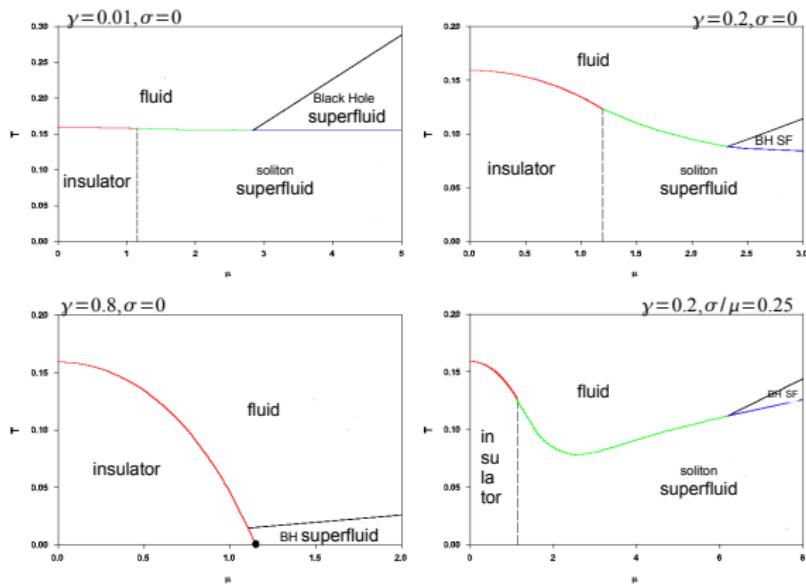
C: soliton → insulator

HC: hairy soliton → superfluid



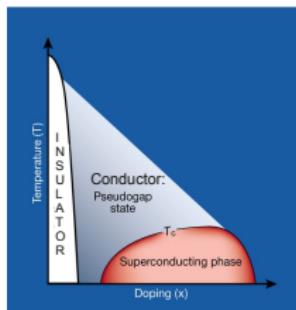
# Holographic superfluids: phase diagrams

(Y. Brihaye & B. Hartmann, Phys. Rev. D83 (2011) 126008)



# Holographic superconductors

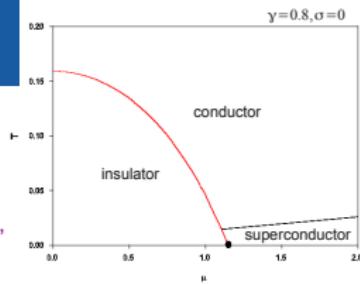
- $\sigma = 0 \rightarrow$  insulator/conductor/superconductor interpretation



(a) Principle phase diagram  
of a cuprate superconductor

(from: <http://www.pha.jhu.edu/~vstanev/>)

(b) taken from  
Brihaye and B. Hartmann,  
Phys. Rev. D, 2011



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# Summary

- Emerging evidence that some condensed matter phenomena are described by **strongly coupled** Quantum Field Theories (QFTs)
- **AdS/CFT** connects strongly coupled QFTs to (classical) gravity theories
- **Holographic superconductors/superfluids ...**
  - ... can be described qualitatively by “hairy” black holes/solitons
  - ... some quantitative results agree with experimental ones:  
 $\omega_g/T_c \approx 8$
- Interesting in both directions
  - deeper insight into “**No hair**” theorems for black holes
  - “AdS/CFT in the laboratory”?



# Holographic superfluids: Free energy

(Y. Brihaye & B. Hartmann, Phys. Rev. D83 (2011) 126008)

- Black strings (BS)

$$\left( \frac{\Omega}{V_2} \right)_{\text{BS}} = \textcolor{orange}{C_l} - 2\textcolor{orange}{C_z}$$

- Solitonic solutions (S)

$$\left( \frac{\Omega}{V_2} \right)_{\text{S}} = \textcolor{orange}{C_l} + \textcolor{orange}{C_z}$$

where

$$f(\rho \gg 1) = \rho^2 + \frac{(\textcolor{orange}{C_l} + \textcolor{orange}{C_z})}{\rho} + O(\rho^{-2}) , \quad b(\rho \gg 1) = \rho^2 + \frac{\textcolor{orange}{C_l}}{\rho} + O(\rho^{-2}) ,$$

$$p(\rho \gg 1) = \rho^2 + \frac{\textcolor{orange}{C_z}}{\rho} + O(\rho^{-2}) , \quad g(\rho \gg 1) \sim O(\rho^{-3})$$