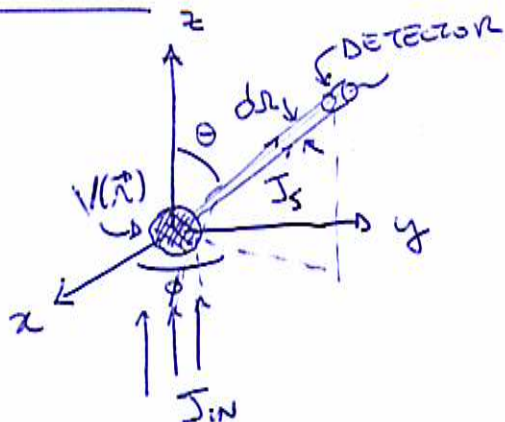


1 - CROSS SECTION



$I_0 =$ INTENSITY OF THE INCIDENT BEAM

$$= \frac{\text{NR. PARTICLES INCIDENT}}{\text{TIME} \times \text{AREA} \perp \hat{z}}$$

$$dm = \frac{\text{NR. PARTICLES HITTING THE DETECTOR}}{\text{TIME}}$$

DEFINITION OF DIFFERENTIAL CROSS SECTION: $dm = \sigma(\theta, \phi) I_0 d\Omega$

OR $\sigma(\theta, \phi) = \frac{1}{I_0} \frac{dm}{d\Omega}$, NOTICE $[\sigma] = \text{AREA}$

IN MANY CASES, THE DIFFERENTIAL CROSS SECTION IS DENOTED BY $\frac{d\sigma(\theta, \phi)}{d\Omega}$

TOTAL CROSS SECTION: $\sigma_{\text{tot}} = \int \sigma(\theta, \phi) d\Omega = \int \sigma(\theta, \phi) \sin\theta d\theta d\phi$

FOR THE NOTATION $\frac{d\sigma}{d\Omega} \Rightarrow \sigma_{\text{tot}} = \int \frac{d\sigma(\theta, \phi)}{d\Omega} d\Omega$

IN TERMS OF THE PROBABILITY FLUXES \vec{J}_{in} AND \vec{J}_s ,

$$\sigma(\theta, \phi) d\Omega = \frac{J_s r^2 d\Omega}{J_{\text{in}}} \rightarrow \sigma = \frac{r^2 J_s}{J_{\text{in}}}$$

WITH $r^2 d\Omega \cong \text{AREA OF THE DETECTOR}$

2 - HYPOTHESIS

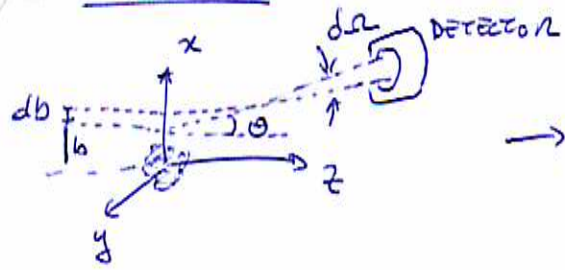
HERE, WE WILL BE INTERESTED IN THE CASE OF

a) LOW-INTENSITY BEAM. ACTUALLY, WE ARE INTERESTED IN THE CASE OF NON-INTERACTING PARTICLES.

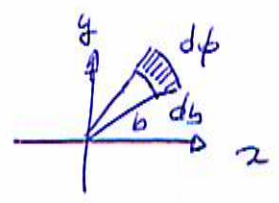
b) ELASTIC SCATTERING

c) FAR-FIELD AND STEADY-STATE REGIME: THE DETECTOR IS FAR FROM THE SCATTERING CENTER AND WE ARE NOT INTERESTED IN NON-TRANSIENT

SSICAT CASE



→ IN THE xy PLANE



$dA = b db dφ$

$dm = I_0 dA = I_0 b db dφ$

$\Rightarrow \sigma = \frac{1}{I_0} \times \frac{I_0 dA}{d\Omega} = \frac{1}{I_0} \times \frac{I_0 b db dφ}{\sin\theta db dφ} \Rightarrow \boxed{\sigma(\theta, \phi) = \frac{b}{\sin\theta} \frac{db}{d\theta}}$

• FOR THE COULOMB POTENTIAL

$V(\vec{r}) = V(r) = \frac{z_1 z_2 e^2}{4\pi\epsilon_0 r}$

$\Rightarrow b(\theta) = \frac{D}{2} \cotg\left(\frac{\theta}{2}\right)$

with $D = \frac{z_1 z_2 e^2}{4\pi\epsilon_0 \frac{mv^2}{2}}$

$\Rightarrow \boxed{\sigma \equiv \sigma(\theta) = \left(\frac{1}{4\pi\epsilon_0}\right)^2 \left(\frac{z_1 z_2 e^2}{2mv^2}\right)^2 \frac{1}{\sin^4(\theta/2)}} = \frac{1}{16} \left(\frac{z_1 z_2 e^2}{4\pi\epsilon_0 E_k}\right) \frac{1}{\sin^4(\theta/2)}$
 $E_k = \frac{1}{2}mv^2 = \frac{1}{2}p^2$

4- QUANTUM TREATMENT

HERE, WE DO NOT HAVE A TRAJECTORY.

INSTEAD, EACH PARTICLE OF THE INCIDENT BEAM IS DESCRIBED AS A WAVEPACKAGE SUCH THAT, AT $t=0$, IT IS "BEHIND" THE SCATTERING POTENTIAL

$\psi(\vec{r}, 0) = \int \varphi(\vec{k}) e^{i\vec{k} \cdot (\vec{r} - \vec{r}_0)} d^3k$

WHERE $\varphi(\vec{k}) \equiv$ FUNCTION CENTERED AT $\vec{k}_0 = (0, 0, k_0)$
 $\vec{r}_0 \equiv$ CENTER OF THE WAVEPACKAGE $\equiv -r_0 \hat{z}$

EXAMPLE: $\varphi(\vec{k}) = \frac{1}{\pi^{3/4} (\alpha\beta\gamma)^{3/2}} \exp\left\{ -\frac{(k_z - k_0)^2}{2\alpha^2} - \frac{k_x^2}{2\beta^2} - \frac{k_y^2}{2\gamma^2} \right\}$

$\Rightarrow \Delta k_z = \alpha, \quad \Delta k_x = \beta, \quad \Delta k_y = \gamma$

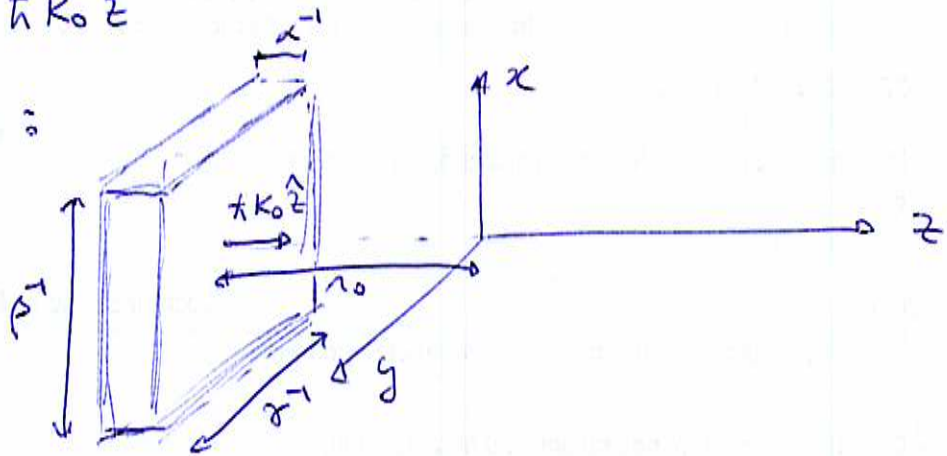
$\Rightarrow \psi(\vec{r}, \omega) = \frac{\sqrt{\alpha\beta\gamma}}{\pi^{3/4}} \exp\left\{ -\frac{\alpha^2}{2} (z+r_0)^2 - \frac{\beta^2}{2} x^2 - \frac{\gamma^2}{2} y^2 + i k_0 (z+r_0) \right\}$

THUS, $\Delta z = \alpha^{-1}$ $\Delta x = \beta^{-1}$ $\Delta y = \gamma^{-1}$

(3)

$\langle \vec{p} \rangle = \hbar k_0 \hat{z}$

• FOR $\frac{\text{small}}{\beta}$ AND γ :



TIME-EVOLUTION = NOW, WE NEED TO EVOLVE $\psi(\vec{r}, 0)$ FOR

$$H = \frac{p^2}{2m} + V(\vec{r})$$

IN THE PREVIOUS EXAMPLE, THEN

$$|\psi(\vec{r}, t)\rangle = U(t, 0) |\psi(\vec{r}, 0)\rangle$$

• FOR THE CASE $V(\vec{r}) = 0 \Rightarrow$ SIMPLE TIME EVOLUTION

$$\psi(\vec{r}, t) = \frac{1}{(2\pi)^{3/2}} \int \psi(\vec{k}) e^{i\vec{k} \cdot (\vec{r} - \vec{r}_0)} e^{-i \frac{E_k t}{\hbar}} d^3k$$

with $E_k = \frac{\hbar^2 k^2}{2m}$

• FOR $V(\vec{r}) \neq 0$, WE WILL ASSUME THAT

a) THE SOLUTIONS $H\psi = E\psi$ EXIST

b) THEY ARE OF TYPE

$$\psi_{\vec{k}}^{(+)}(\vec{r}) = \frac{1}{(2\pi)^{3/2}} \left(e^{i\vec{k} \cdot \vec{r}} + f_{\vec{k}}(\theta, \phi) \frac{e^{i\vec{k} \cdot \vec{r}}}{r} \right)$$

\downarrow
 SCATTERING AMPLITUDE

$$= \psi_{IN} + \psi_S$$

c) $\psi(\vec{r}, 0) = \int \psi(\vec{k}) e^{-i\vec{k} \cdot \vec{r}_0} \psi_{\vec{k}}^{(+)}(\vec{r}) d^3k$

$$\psi(\vec{r}, t) = \int \psi(\vec{k}) e^{-i\vec{k} \cdot \vec{r}_0} \psi_{\vec{k}}^{(+)}(\vec{r}) e^{-i \frac{E_k t}{\hbar}} d^3k$$

with $E_k = \frac{\hbar^2 k^2}{2m}$ FAR FROM THE SCATTERING POTENTIAL

HYPOTHESIS (b) IS VERY PLAUSIBLE. THE SCATTERED (4)

WAVE $\psi_s \propto f(\theta, \phi) \frac{e^{ikr}}{r}$ IS JUST AN OUTGOING

SPHERICAL WAVE WEIGHTED BY $f(\theta, \phi)$ - THIS SHOULD BE OKAY FOR A FARAWAY DETECTOR.

EXAMPLE: ID: $\psi_k^+ = \psi_{in} + \psi_s = \begin{cases} A e^{ikx} + B e^{-ikx} & , x \rightarrow -\infty \\ C e^{ikx} & , x \rightarrow \infty \end{cases}$

where $\psi_{in} = A e^{ikx}$

• CONNECTION BETWEEN THE DIFFERENTIAL CROSS SECTION AND THE SCATTERING AMPLITUDE

INCIDENT FLUX: $\vec{J}_{in} = \frac{\hbar}{m} \text{Im}(\psi_{in}^* \nabla \psi_{in})$, $\psi_{in} = A_k e^{i\vec{k} \cdot \vec{r}} = A_k e^{ikz}$

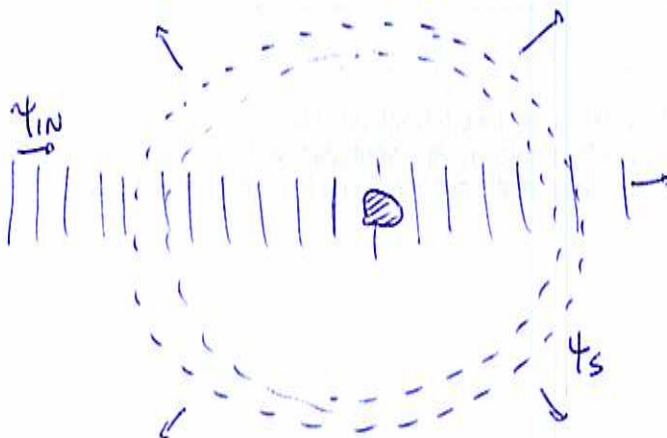
$$\vec{J}_{in} = |A_k|^2 \frac{\hbar}{m} \vec{k}, \quad \vec{k} = k \hat{z}$$

SCATTERED FLUX: $\vec{J}_s = \frac{\hbar}{m} \text{Im}(\psi_s^* \nabla \psi_s)$, $\psi_s = A_k f(\theta, \phi) \frac{e^{ikr}}{r}$

$$\nabla \psi_s = \hat{r} \frac{\partial}{\partial r} \psi_s + \hat{\theta} \frac{1}{r} \frac{\partial}{\partial \theta} \psi_s + \hat{\phi} \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} \psi_s = A_k \left(f(\theta, \phi) ik \frac{e^{ikr}}{r} \hat{r} + O(r^{-2}) \right)$$

$$\Rightarrow \vec{J}_s \approx |A_k|^2 |f_k(\theta, \phi)|^2 \frac{\hbar k}{m} \frac{\hat{r}}{r^2}$$

$$\Rightarrow \sigma(\theta, \phi) = |f_k(\theta, \phi)|^2$$



ALTERNATIVELY, WE COULD ASK ABOUT THE PROBABILITY OF FINDING THE PARTICLE AT THE DETECTOR BETWEEN t AND $t+dt$ (5)

$$P(t) = |\psi|^2 dV = |A_k|^2 |f_k(\theta, \phi)|^2 \times \frac{1}{r^2} dV, \quad dV = r^2 d\Omega dr$$

$$= |A_k|^2 |f_k(\theta, \phi)|^2 v dt d\Omega$$

with $dr = v dt$

$$dm = \frac{\text{NR. OF PARTICLES}}{\text{TIME}} \times P(t)$$

$$I_0 = \frac{\text{NR. OF PARTICLES}}{\text{TIME}} \times |\psi|^2 \frac{dV}{dA} = \frac{\text{NR. PARTICLES}}{\text{TIME}} \times |A_k|^2 dz, \quad \text{BUT } dz = v dt$$

$$\text{AGAIN} \quad \sigma(\theta, \phi) = \frac{1}{I_0} \frac{dm}{d\Omega} = |f(\theta, \phi)|^2$$

SOLUTION OF THE TIME-INDEPENDENT SCHRÖDINGER EQUATION

$$H\psi = E\psi \rightarrow \left(-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right) \psi = E\psi$$

$$\text{DEFINING } k^2 = \frac{2mE}{\hbar^2} \quad \text{AND } U = \frac{2mV}{\hbar^2}$$

$$\Rightarrow \boxed{(\nabla^2 + k^2)\psi = U\psi}$$

THE GENERAL SOLUTION IS OBTAINED VIA GREEN'S FUNCTION

$$\psi_{\vec{k}}(\vec{r}) = \psi_{\vec{k}}^{(0)}(\vec{r}) - \frac{1}{4\pi} \int G(\vec{r}, \vec{r}') U(\vec{r}') \psi_{\vec{k}}(\vec{r}') d^3r'$$

$$\text{WHERE } (\nabla^2 + k^2) G(\vec{r}, \vec{r}') = -4\pi \delta(\vec{r} - \vec{r}') \quad \text{AND } (\nabla^2 + k^2) \psi^{(0)} = 0$$

$$\text{PROOF: } \nabla^2 \psi = \nabla^2 \psi^{(0)} - \frac{1}{4\pi} \int (\nabla^2 G) U(\vec{r}') \psi(\vec{r}') d^3r'$$

$$= \nabla^2 \psi^{(0)} - \frac{1}{4\pi} \int (-4\pi \delta(\vec{r} - \vec{r}') - k^2 G) U \psi d^3r'$$

$$= -k^2 \psi^{(0)} + U(\vec{r}) \psi(\vec{r}) + k^2 \left[\frac{1}{4\pi} \int G U \psi d^3r' \right]$$

THUS,

$$\nabla^2 \psi + k^2 \psi = U \psi$$

(6)

COMPUTING G

a) ~~ASSUME~~ NOTICE $G(\vec{n}, \vec{n}) = G(\vec{n} - \vec{n})$, BECAUSE

$\nabla^2 + k^2$ AND $\delta(\vec{n} - \vec{n})$ ARE INVARIANT UNDER TRANSLATIONS OF \vec{n} AND \vec{n}

b) FOURIER TRANSFORM: $G(\vec{n}) = \int g(\vec{k}) e^{i\vec{k} \cdot \vec{n}} d\vec{k}$

$$\delta(\vec{n}) = \frac{1}{(2\pi)^3} \int e^{i\vec{k} \cdot \vec{n}} d\vec{k}$$

c) SOLVE FOR THE FREE PARTICLE: $(\nabla^2 + k^2) G(\vec{n} - \vec{n}) = -4\pi \delta(\vec{n} - \vec{n})$

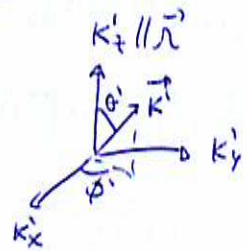
$$\Rightarrow \int (\nabla^2 + k^2) e^{i\vec{k} \cdot \vec{n}} g(\vec{k}) d\vec{k} = -\frac{4\pi}{(2\pi)^3} \int e^{i\vec{k} \cdot \vec{n}} d\vec{k}$$

$$\Rightarrow \int (k^2 - k'^2) e^{i\vec{k}' \cdot \vec{n}} g(\vec{k}') d\vec{k}' = -\frac{1}{2\pi^2} \int e^{i\vec{k}' \cdot \vec{n}} d\vec{k}'$$

$$\Rightarrow g(\vec{k}') = \frac{1}{2\pi^2} \times \frac{1}{k'^2 - k^2}$$

$$\Rightarrow G(\vec{n}) = \frac{1}{2\pi^2} \int \frac{e^{i\vec{k}' \cdot \vec{n}}}{k'^2 - k^2} d\vec{k}'$$

INTEGRATING OVER \vec{k}' : SET $k'_z \parallel \vec{n}$



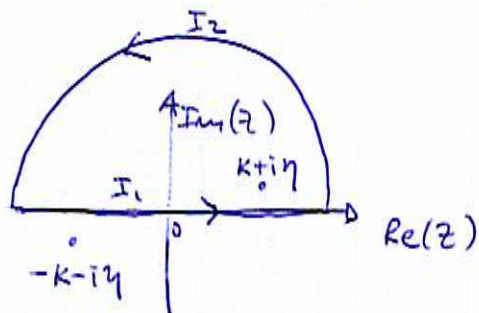
$$G(\vec{n}) = \frac{2\pi}{2\pi^2} \int \frac{e^{i k'_z n \cos \theta}}{k'^2 - k^2} k'^2 \sin \theta' dk' d\theta'$$

$$\begin{aligned} & \text{(\cos \theta' = \mu)} \\ & \text{(-\sin \theta' d\theta' = d\mu)} \end{aligned} \quad \frac{1}{\pi} \int_0^1 \int_{-1}^1 \frac{e^{i k'_z n \mu}}{k'^2 - k^2} d\mu k'^2 dk' = i \frac{1}{\pi n} \int_0^\infty \frac{e^{i k' n} - e^{-i k' n}}{k' (k'^2 - k^2)} k'^2 dk'$$

$$= -\frac{1}{\pi n} \frac{d}{dn} \left(\int_{-\infty}^\infty \frac{e^{i k' n}}{k'^2 - k^2} dk' \right)$$

COMPUTING THE INTEGRAL VIA RESIDUES

$$I = \int_{-\infty}^{\infty} \frac{e^{ik'l} dk'}{[k' - (k+i\eta)][k' + (k+i\eta)]} \quad \text{with } \eta \rightarrow 0^+$$



$$I_1 + I_2 = 2\pi i \sum R(z_i)$$

$$I_1 = \int_{-\infty}^{\infty} \frac{e^{ik'l}}{k^2 - k^2} dk' = I$$

$$I_2 = \int_0^\pi \frac{e^{i|z|(\cos\theta + i\sin\theta)l}}{e^{2i\theta} z^2 - k^2} |z| d\theta \quad \begin{matrix} |z| \rightarrow \infty \\ \rightarrow 0 \end{matrix}$$

$$\Rightarrow I_1 = 2\pi i * \frac{e^{i(k+i\eta)l}}{2(k+i\eta)} \rightarrow \pi i \frac{e^{ikl}}{k}$$

FINALLY, $G(\vec{r}) = -\frac{1}{\pi\eta} \frac{d}{d\eta} \left(\pi i \frac{e^{ikl}}{k} \right) = \frac{e^{ikl}}{\eta}$

USUALLY, THIS GREEN'S FUNCTION IS DENOTED BY

$$G^{(+)} = \frac{e^{ikl}}{\eta} \rightarrow \text{OUTGOING WAVE}$$

IF WE HAD CHOSEN $\eta \rightarrow 0^-$, THEN THE POLE WOULD HAPPEN FOR $-k$

AND WE WOULD HAVE $G^{(-)} = \frac{e^{-ikl}}{\eta} \rightarrow \text{INCOMING WAVE}$

~~FOR LARGE DISTANCES...~~

BACK TO ψ : $\psi_{\vec{r}}^{(+)}(\vec{r}) = \frac{1}{(2\pi)^{3/2}} \left(e^{i\vec{k}\cdot\vec{r}} - \frac{(2\pi)^{3/2}}{4\pi} \int \frac{e^{ik|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|} U(\vec{r}') \psi_{\vec{r}'}(\vec{r}') d\vec{r}' \right)$

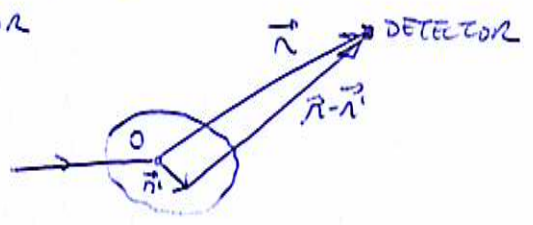
FOR $\eta \rightarrow 0$, ONE COULD THINK IN DOING BETTER JOB: EXPANDING TO $\eta \gg \eta'$ $\frac{e^{ik|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|} \approx \frac{e^{ikl}}{\eta}$, BUT THIS YIELDS ψ_S WITH NO DEPENDENCE ON θ AND ϕ .

$|\vec{r}-\vec{r}'| = \sqrt{\eta^2 - 2\vec{r}\cdot\vec{r}' + \eta'^2} \approx \eta \left(1 - \frac{\vec{r}\cdot\vec{r}'}{\eta^2} \right) = \eta - \vec{r}\cdot\hat{n}$

Thus, $\frac{1}{|\vec{n}-\vec{n}'|} \approx \frac{1}{r} \left(1 + \frac{\vec{n}' \cdot \hat{n}}{r}\right) \rightarrow \frac{1}{r}$

BECAUSE IT INTRODUCES ERRORS THAT VANISH FOR $r \rightarrow \infty$. THIS IS NOT SO FOR THE EXPONENTIAL FACTOR

$e^{iK|\vec{n}-\vec{n}'|} \approx e^{iKr} \times e^{-iK\hat{n} \cdot \vec{n}'}$



BUT $K\hat{n}$ IS THE SCATTERED WAVE VECTOR, THUS, LET'S CALL $\vec{k}_s = K\hat{n}$

FINALLY

$\psi_{\vec{k}}^{(+)}(\vec{r}) \approx \frac{e^{i\vec{k} \cdot \vec{r}}}{(2\pi)^{3/2}} - \frac{1}{4\pi} \int \frac{e^{i\vec{k}\vec{n}}}{r} e^{-i\vec{k}_s \cdot \vec{n}'} U(\vec{n}') \psi_{\vec{k}}^{(+)}(\vec{n}') d\vec{n}'$

THUS

$f_{\vec{k}}(\theta, \phi) = f_{\vec{k}}(\hat{n}) = -\frac{(2\pi)^{3/2}}{4\pi} \int e^{-i\vec{k}\hat{n} \cdot \vec{n}'} U(\vec{n}') \psi_{\vec{k}}^{(+)}(\vec{n}') d\vec{n}'$

BORN SERIES (OR BORN APPROXIMATION)

SO FAR, $\psi_{\vec{k}} = \psi_{IN} + \psi_S = \psi_{\vec{k}}^{(0)} - \frac{1}{4\pi} \int G(\vec{n}-\vec{n}') U(\vec{n}') \psi_{\vec{k}}(\vec{n}') d\vec{n}'$

IN KET FORM,

$\langle \vec{n} | \psi_{\vec{k}} \rangle = \langle \vec{n} | \psi_{\vec{k}}^{(0)} \rangle - \frac{1}{4\pi} \int \langle \vec{n} | G | \vec{n}' \rangle \langle \vec{n}' | U | \psi_{\vec{k}} \rangle d\vec{n}'$
 $= \langle \vec{n} | \psi_{\vec{k}}^{(0)} \rangle - \frac{1}{4\pi} \langle \vec{n} | G U | \psi_{\vec{k}} \rangle$

$\Rightarrow |\psi_{\vec{k}} \rangle = |\psi_{\vec{k}}^{(0)} \rangle - \frac{1}{4\pi} G U |\psi_{\vec{k}} \rangle$

$\Rightarrow |\psi_{\vec{k}} \rangle = \frac{1}{1 + \frac{1}{4\pi} G U} |\psi_{\vec{k}}^{(0)} \rangle$

RECALLING THAT

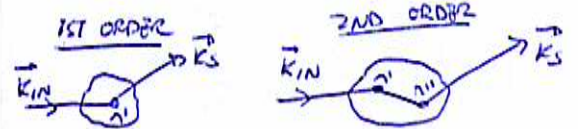
$$\frac{1}{1-x} = \sum_{n=0}^{\infty} x^n$$

$$\Rightarrow |\Psi_{\vec{k}}\rangle = \left(1 - \frac{G U}{4\pi} + \frac{1}{4\pi} G U G U - \dots \right) |\Psi_{\vec{k}}^{(0)}\rangle$$

IN THE SPACE REPRESENTATION

$$\Psi_{\vec{k}}(\vec{r}) = \Psi_{\vec{k}}^{(0)}(\vec{r}) - \frac{1}{4\pi} \int G(\vec{r}-\vec{r}') U(\vec{r}') \Psi_{\vec{k}}^{(0)}(\vec{r}') d\vec{r}' + \frac{1}{(4\pi)^2} \int \int G(\vec{r}-\vec{r}') U(\vec{r}') G(\vec{r}'-\vec{r}'') U(\vec{r}'') \Psi_{\vec{k}}^{(0)}(\vec{r}'') d\vec{r}' d\vec{r}'' + \dots$$

ALTERNATIVELY, WE COULD START FROM



$$\Psi_{\vec{k}} = \Psi_{\vec{k}}^{(0)} - \frac{1}{4\pi} \int G U \Psi_{\vec{k}}$$

AND DOWN THE APPROXIMATION THAT $\Psi_{\vec{k}} \approx \Psi_{\vec{k}}^{(0)}$ IN THE INTEGRAL. THEN

$$\Psi_{\vec{k}} \approx \Psi_{\vec{k}}^{(0)} - \frac{1}{4\pi} \int G U \Psi_{\vec{k}}^{(0)}$$

WHICH IS THE BORN APPROXIMATION IN 1ST ORDER IN U

IF WE WISH THE SECOND ORDER APPROXIMATION, THEN WE FEED THE 1ST-ORDER ONE ON THE ORIGINAL EQUATION, YIELDING

$$\Psi_{\vec{k}} \approx \Psi_{\vec{k}}^{(0)} - \frac{1}{4\pi} \int G U \Psi_{\vec{k}}^{(0)} + \frac{1}{(4\pi)^2} \int \int G U G U \Psi_{\vec{k}}^{(0)}$$

NOTICE $f(\vec{k}_{IN}, \vec{k}_S) \equiv f(\vec{k}_{IN} - \vec{k}_S)$
MINIMUM TRANSFER

REPEATING THIS PROCEDURE WE RECOVER THE BORN SERIES

TO 1ST ORDER

$$f_{\vec{k}}(\vec{r}) = f_{\vec{k}_{IN}}(\vec{k}_S) \approx -\frac{1}{4\pi} \int e^{i(\vec{k}_{IN} - \vec{k}_S) \cdot \vec{r}} U(\vec{r}) d\vec{r}$$

WHICH IS THE FOURIER TRANSFORM OF THE POTENTIAL

VALIDITY OF THE BORN APPROXIMATION

UP TO 1ST ORDER IN V ,

$$\Psi_{\vec{k}}(\vec{r}) = \Psi_{\vec{k}}^{(0)}(\vec{r}) - \frac{1}{4\pi} \int G(\vec{r}-\vec{r}') U(\vec{r}') \Psi_{\vec{k}}^{(0)}(\vec{r}') d\vec{r}'$$

↓
T MATRIX
PAGE 9.2

MEANING THAT WE SUBSTITUTED $\Psi = \Psi_{IN} + \Psi_S$ BY Ψ_{IN} IN THE INTEGRAND

THUS, IN A ROUGH SENSE, WE COULD SAY THAT

$$|\Psi_S| \ll |\Psi_{\vec{k}}^{(0)}| \quad \text{IN THE REGION OF THE POTENTIAL}$$

Thus, $f(\vec{k}_{in}, \vec{k}_s) = -\frac{1}{4\pi} \times (2\pi)^3 \times \frac{2m}{\hbar^2} \int \frac{e^{-i\vec{k}_s \cdot \vec{r}}}{(2\pi)^{3/2}} V(\vec{r}) \frac{e^{i\vec{k}_{in} \cdot \vec{r}}}{(2\pi)^{3/2}} d\vec{r}$ (9a)

$$\langle \vec{k}_s | \vec{r} \rangle \langle \vec{r} | V | \vec{k}_{in} \rangle$$

$$\Rightarrow f(\vec{k}_{in}, \vec{k}_s) = -\frac{1}{4\pi} (2\pi)^3 \frac{2m}{\hbar^2} \langle \vec{k}_s | T^{(1)} | \vec{k}_{in} \rangle$$

With $T^{(1)} = V$ BEING THE TRANSFER MATRIX OPERATOR (OR THE TRANSITION OPERATOR) IN 1ST ORDER IN V

USING ALL ORDERS, THEN

$$T = V + V G U + V G U G U + V G U G U G U + \dots$$

IT IS POSSIBLE TO SHOW THAT THIS IS EQUAL TO

$$T = V + V \frac{1}{E - H_0 + i\eta} V + V \frac{1}{E - H_0 + i\eta} V \frac{1}{E - H_0 + i\eta} V + \dots$$

$$= \frac{V}{1 - \frac{1}{E - H_0 + i\eta} V} = V + V \frac{1}{E - H_0 + i\eta} T$$

ROUGHLY, THE "DISTORTION" OF THE INCIDENT WAVE IS SMALL $\frac{|\psi_S(\omega)|}{|\psi_{\vec{k}}^{(0)}(\omega)|} \ll 1 \Rightarrow \left| \int e^{i(\vec{k}_{in} - \vec{k}_s) \cdot \vec{r}} \frac{U(\vec{r})}{r} d\vec{r} \right| \ll 1$ (10)

with $\vec{k}_{in} = k \hat{z}$ AND $\vec{k}_s = k \hat{n}$, CHOOSING $\hat{z} // \hat{z}$

THEN, $\left| \int e^{ikr} e^{-ikr' \cos \theta'} \frac{U(\vec{r}')}{r'} d\vec{r}' \right| \ll 1$ CHOOSING $U(\vec{r}') = U(r')$

$\Rightarrow 2\pi \int e^{ikr'} \frac{e^{-ikr'} - e^{ikr'}}{-ikr'} \frac{U(r')}{r'} r'^2 dr' = \frac{4\pi}{k} \int e^{ikr'} \sin kr' U(r') dr'$

THUS, $\frac{2m}{\hbar^2 k} \left| \int e^{ikr'} \sin(kr') V(r') dr' \right| \ll 1$

• LOW-ENERGY ANALYSIS: $(kr' \rightarrow 0) \Rightarrow e^{ikr'} \sim kr' \rightarrow kr'$

$\Rightarrow \frac{2m}{\hbar^2} \left| \int r' V(r') dr' \right| \ll 1 \Rightarrow \frac{2m}{\hbar^2} \frac{|V_0| R_0^2}{2} \ll 1$

(MAKE $V=V_0$ AND OF SIZE R_0)

THUS, $V_0 \ll \frac{\hbar^2}{m R_0^2}$

PHYSICAL MEANING $\rightarrow V_0 \ll \frac{\hbar^2 k^2}{2m}$, $k = \frac{\sqrt{2mE}}{\hbar} = \frac{\sqrt{2}}{R_0}$

WHICH IS THE WAVE NUMBER OF A PARTICLE CONFINED IN A REGION OF ORDER R_0

\Rightarrow THE POTENTIAL HAS TO BE TOO SHALLOW IN ORDER TO BIND A PARTICLE

• HIGH-ENERGY ANALYSIS: $(kR_0 \gg 1) \Rightarrow e^{ikr'} \sim -\frac{1}{2i}$
 SINCE $e^{2ikr'}$ OSCILLATES RAPIDLY

$\Rightarrow \frac{2m}{\hbar^2 k} \cdot \frac{1}{2} V_0 R_0 \ll 1 \Rightarrow \boxed{V_0 \ll \frac{\hbar^2}{m R_0^2} * (k R_0)}$

THUS, IF THE BORN APPROXIMATION WORKS IN THE LOW-ENERGY REGIME, IT ALSO WORKS IN THE HIGH-ENERGY ONE.

EXAMPLES:

(11)

1) YUKAWA POTENTIAL: $V(\vec{r}) = V(r) = \frac{V_0 e^{-\alpha r}}{\alpha r}$, $\alpha > 0$

2) COULOMB POTENTIAL: $V(r) = \lim_{\alpha \rightarrow \infty} \frac{V_0 e^{-\alpha r}}{\alpha r} = \left(\frac{V_0}{\alpha}\right) \frac{e^{-\alpha r}}{r} \rightarrow \frac{V_0}{\alpha r}$

3) SPHERICAL BARRIER: $V(\vec{r}) = V(r) = \begin{cases} V_0, & \text{if } r \leq R \\ 0, & \text{if } r > R \end{cases}$

BORN APPROXIMATION IN 1st-ORDER TO SPHERICAL POTENTIALS

$$f_{\vec{k}_i, \vec{k}_s}(\vec{k}_s) = \frac{-1}{4\pi} \int e^{i(\vec{k}_i - \vec{k}_s) \cdot \vec{r}} \frac{2m}{\hbar^2} V(\vec{r}) d\vec{r}$$

LET US TAKE $\hat{z} \parallel \Delta\vec{k} = \vec{k}_i - \vec{k}_s$

$$= \frac{-2\pi}{4\pi} + \frac{2m}{\hbar^2} \iint e^{i\Delta k r \cos\theta} V(r) r^2 \sin\theta dr d\theta$$

$$f_{\vec{k}_i, \vec{k}_s}(\vec{k}_s) = \frac{-2m}{\Delta k \hbar^2} \int_0^\infty \sin(\Delta k r) V(r) r dr$$

NOTICE $f_{\vec{k}_i, \vec{k}_s}(\vec{k}_s) = f(\theta)$ AND DOES NOT DEPEND ON ϕ

AND $f(\theta) \in \mathbb{R}$



$$\begin{aligned} \vec{k}_i &= k \hat{z} \\ \vec{k}_s &= k \hat{n} \end{aligned}$$

$$\Delta k = \sqrt{k_i^2 - 2\vec{k}_i \cdot \vec{k}_s + k_s^2} = k \sqrt{2 - 2\cos\theta} = k \sqrt{4\sin^2 \frac{\theta}{2}} = 2k \left| \sin \frac{\theta}{2} \right|$$

1) YUKAWA POTENTIAL

$$f(\theta) = \frac{-2m}{\hbar^2 2k \left| \sin \frac{\theta}{2} \right|} \int_0^\infty x \sin(2k \left| \sin \frac{\theta}{2} \right| x) \frac{V_0}{\alpha x} e^{-\alpha x} dx$$

$$= \frac{-2m}{\hbar^2 \Delta k} \left(\frac{V_0}{\alpha}\right) \times \frac{1}{2i} \int_0^\infty \left(e^{i(\Delta k - \alpha)x} - e^{-i(\Delta k + \alpha)x} \right) dx$$

$$= \frac{-2m}{\hbar^2 \Delta k} \left(\frac{V_0}{\alpha}\right) \times \frac{1}{2i} \left[\frac{1}{\alpha - i\Delta k} - \frac{1}{\alpha + i\Delta k} \right] = \frac{-2m}{\hbar^2 \Delta k} \frac{V_0}{\alpha} \frac{\Delta k}{\alpha^2 + \Delta k^2}$$

FINALLY , $f(\theta) = -\frac{2m}{\hbar^2} \left(\frac{V_0}{\alpha} \right) \frac{1}{\alpha^2 + \Delta k^2}$

(12)

$$\Rightarrow \left(\sigma = \sigma(\theta) = \frac{4m^2}{\hbar^4} \left(\frac{V_0}{\alpha} \right)^2 \left(\frac{1}{\alpha^2 + 4k^2 \sin^2(\theta/2)} \right)^2 \right)^2$$

→ PROBABLY MISLEADING, BECAUSE THE BORN APPROXIMATION IS NOT GOOD FOR LOW ENERGIES

2) COULOMB POTENTIAL , TAKING $\alpha \rightarrow 0$ AND $\frac{V_0}{\alpha} \rightarrow \frac{ZZe^2}{4\pi\epsilon_0}$

$$\Rightarrow \sigma(\theta) = \frac{4m^2}{\hbar^4} \left(\frac{ZZe^2}{4\pi\epsilon_0} \right)^2 \frac{1}{(4k^2)^2 \sin^4(\theta/2)}$$

IDENTIFYING $\frac{\hbar^2 k^2}{2m} = E_k$

AS THE KINETIC ENERGY OF THE INCIDENT PARTICLE BEAM

$$\Rightarrow \left(\sigma(\theta) = \frac{1}{16} \left(\frac{ZZe^2}{4\pi\epsilon_0 E_k} \right)^2 \frac{1}{\sin^4(\theta/2)} \right)$$

WHICH IS EXACTLY THE CLASSICAL RESULT

$$\begin{aligned} 3) f(\theta) &= \frac{-2m}{\hbar^2 \Delta k} \int_0^a x \sin(\Delta k x) V_0 dx = \frac{-2mV_0}{\hbar^2 \Delta k} \left[-\frac{x \cos(\Delta k x)}{\Delta k} \right]_0^a + \frac{1}{\Delta k} \int_0^a \cos(\Delta k x) dx \\ &= \frac{-2mV_0}{\hbar^2 \Delta k} \left[\frac{-a \cos(\Delta k a)}{\Delta k} + \frac{1}{\Delta k^2} \sin(\Delta k a) \right] \end{aligned}$$

$$\Rightarrow \left(f(\theta) = \frac{2mV_0}{\hbar^2 \Delta k^3} \left[\Delta k a \cos(\Delta k a) - \sin(\Delta k a) \right] \right)$$

$$\sigma(\theta) = |f(\theta)|^2$$

METHOD OF PARTIAL WAVES

- SCATTERING BY A CENTRAL POTENTIAL $V(\vec{r}) = V(r)$
- $V(r)$ IS LOCALIZED $\Rightarrow V(r) \rightarrow 0$ FOR $r \rightarrow \infty$

IDEA OF THE METHOD

$$H|\psi\rangle = (H_0 + V(r))|\psi\rangle = E|\psi\rangle$$

- FOR $V=0 \Rightarrow |\psi\rangle = |\vec{k}\rangle = \sum_{k, l, m} a_{k, l, m} |k, l, m\rangle$

Where $|k, l, m\rangle = |k\rangle \otimes |l, m\rangle$

SPHERICAL HARMONICS

$$\langle \vec{r} | k, l, m \rangle = R_{k, l}(r) Y_{l, m}(\theta, \phi)$$

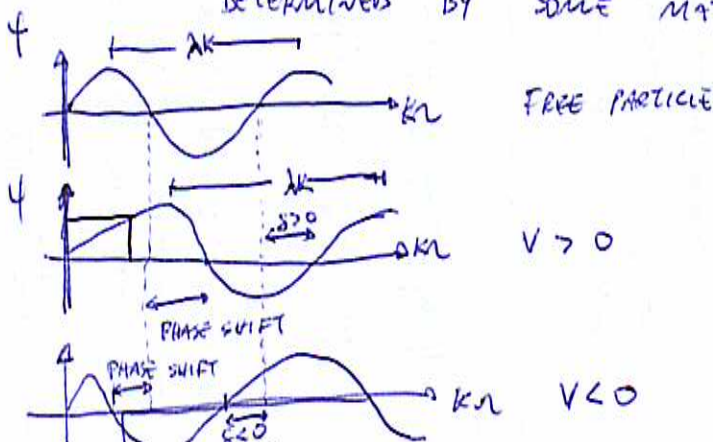
\Rightarrow PLANE WAVE = SUPERPOSITION OF FREE SPHERICAL WAVES
 = SUPERPOSITION OF PARTIAL WAVES

- FOR $V \neq 0 \Rightarrow |\psi\rangle = \sum_{k, l, m} a_{k, l, m} |k, l, m\rangle$

Where $\langle \vec{r} | k, l, m \rangle = R_{k, l}(r) Y_{l, m}(\theta, \phi)$

$V(r)$ INFLUENCES ONLY $R_{k, l}(r)$, THE RADIAL PART, NOT THE ANGULAR.

IN ADDITION, FOR $\lambda \rightarrow \infty$, THE FREE-PARTICLE WAVE SHOULD BE RECOVERED BUT WITH SOME PHASE SHIFT WHICH IS DETERMINED BY SOME MATCHING CONSTRAINT



PLANE WAVES IN TERMS OF FREE SPHERICAL WAVES

(14)

(THE CASE OF $V=0$)

SOLVING $\hat{H}_0|\psi\rangle = E|\psi\rangle$, $\hat{H}_0 = \frac{p^2}{2m}$

$\psi_{k\ell m}^{(0)} = R_{k\ell}^{(0)}(r) Y_{\ell m}(\theta, \phi) = \frac{\mu_{k\ell}(r)}{r} Y_{\ell m}(\theta, \phi)$

$\Rightarrow \left(\frac{-\hbar^2}{2m} \frac{d^2}{dr^2} + \frac{\hbar^2 \ell(\ell+1)}{2m r^2} \right) \mu_{k\ell}(r) = E \mu_{k\ell}(r) = \frac{\hbar^2 k^2}{2m} \mu_{k\ell}(r)$

↓
CENTRIFUGAL POTENTIAL

GENERAL SOLUTION: $R_{k\ell}^{(0)} = \sqrt{\frac{2k^2}{\pi}} \left(A_{\ell} j_{\ell}(kr) + B_{\ell} n_{\ell}(kr) \right)$
 ↓ Spherical Bessel ↓ Spherical Neumann

NORMALIZATION: $|A_{\ell}|^2 + |B_{\ell}|^2 = 1$

$j_{\ell}(z) = (-1)^{\ell} z^{\ell} \left(\frac{1}{z} \frac{d}{dz} \right)^{\ell} \left(\frac{\sin z}{z} \right)$

$n_{\ell}(z) = (-1)^{\ell+1} z^{\ell} \left(\frac{1}{z} \frac{d}{dz} \right)^{\ell} \left(\frac{\cos z}{z} \right)$

ex: $j_0 = \frac{\sin z}{z}$, $j_1 = -\frac{\cos z}{z} + \frac{\sin z}{z^2}$

$n_0 = -\frac{\cos z}{z}$, $n_1 = -\frac{\sin z}{z} + \frac{\cos z}{z^2}$

ASYMPTOTIC BEHAVIOR ($r \rightarrow \infty$)

$j_{\ell}(z) = (-1)^{\ell} z^{\ell} \left(\frac{1}{z} \frac{d}{dz} \right)^{\ell-1} \left(\frac{+\cos z}{z^2} + \frac{\sin z}{z^3} \right) = (-1)^{\ell} z^{\ell} \left(\frac{1}{z} \frac{d}{dz} \right)^{\ell-1} \left(\frac{-\sin z}{z^3} - \frac{\cos z}{z^4} \right)$
 $= (-1)^{\ell} z^{\ell} \times \frac{1}{z^{\ell}} \times \frac{1}{z} \left(\frac{d^{\ell}}{dz^{\ell}} \sin z \right) = \frac{(-1)^{\ell}}{z} \frac{d^{\ell}}{dz^{\ell}} \left(\frac{e^{iz} - e^{-iz}}{2i} \right) = \frac{(-1)^{\ell}}{z} (i)^{\ell} \left(\frac{e^{iz}}{2i} - \frac{(-1)^{\ell} e^{-iz}}{2i} \right)$
 $= \frac{1}{z} \left(\frac{(-i)^{\ell} e^{iz} - i^{\ell} e^{-iz}}{2i} \right) = \frac{1}{z} \left(\frac{e^{-\frac{i\pi}{2}\ell} e^{iz} - e^{\frac{i\pi}{2}\ell} e^{-iz}}{2i} \right)$

$|x_{\ell}(z)| \rightarrow \pm \sin(z - \frac{l\pi}{2})$

ANALOGOUSLY,

$$\left[\psi_e \frac{z \rightarrow \infty}{r} = \frac{\cos(z - \frac{\pi}{2}l)}{z} \right]$$

(15)

FINALLY, BECAUSE $\psi_{l,m}$ DOES NOT DIVERGE AT $r \rightarrow \infty$

$\Rightarrow B_l = 0$ (NO ~~FREE~~ SPHERICAL NEUMANN)

THUS THE BASIS IS $\{ \psi_{l,m} \}$

$$\psi_{l,m}^{(0)}(r, \theta, \phi) = \sqrt{\frac{2k^2}{\pi}} j_l(kr) Y_{l,m}(\theta, \phi)$$

• SOME PHYSICAL PROPERTIES OF THE FREE SPHERICAL WAVES

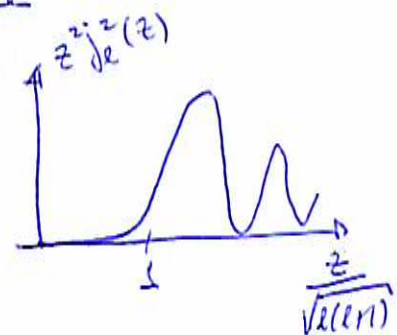
a) WELL DEFINED ANGULAR MOMENTUM

b) BEHAVIOR NEAR THE ORIGIN

$$P(\vec{r}) d\vec{r} \propto j_l^2(kr) |Y_{l,m}(\theta, \phi)|^2 r^2 dr d\Omega$$

FOR SMALL r , $j_l(kr) \sim \frac{(kr)^l}{(2l+1)!!}$

$\Rightarrow P(\vec{r}) \propto r^{2l+2} = r^{2(l+1)}$



IN FACT $z^2 j_l^2(z)$ IS SMALL FOR $z < \sqrt{l(l+1)}$

\Rightarrow PROB. IS SMALL FOR $kr < \sqrt{l(l+1)}$

CONSEQUENCE:

A PARTICLE IN STATE $\psi_{l,m}$ IS WEAKLY AFFECTED BY A POTENTIAL OF ~~SMALL~~ WIDTH

$$b_l(k) = \frac{1}{k} \sqrt{l(l+1)}$$

IN CLASSICAL MECHANICS, THE IMPACT PARAMETER IS OBTAINED FROM $L = m v b$

$\Rightarrow b = \frac{L}{p} \Rightarrow$ IF THE POTENTIAL WIDTH IS $a < b \Rightarrow$ NOTHING HAPPENS \Rightarrow MOTIVE $b_0(k) = \frac{1}{k} \sqrt{l(l+1)} = \underline{b}$

c) ASYMPTOTIC BEHAVIOR FOR $r \rightarrow \infty$

$$\psi_{k\ell m}^{(in)} \xrightarrow{r \rightarrow \infty} + \sqrt{\frac{2k^2}{\pi}} Y_{\ell m}(\theta, \phi) \frac{e^{+i(kr - \frac{2\ell\pi}{2})} - e^{-i(kr - \frac{2\ell\pi}{2})}}{2i k r}$$

$$\propto \frac{Y_{\ell m}(\theta, \phi) \frac{e^{ikr} - e^{-i(kr - \ell\pi)}}{2i k r}}$$

- WHICH IS A SUPERPOSITION OF AN INCOMING WAVE $\frac{e^{-ikr}}{r}$ WITH AN OUTGOING ONE $\frac{e^{ikr}}{r}$.

- SAME AMPLITUDE
- PHASE DIFFERENCE = $\pi\ell$

IF WE CONSTRUCT A ~~WAVE~~ SPHERICAL WAVE PACKET, LOCATED AT $r \rightarrow \infty$ AT $t \rightarrow -\infty$, AND LET IT EVOLVE, THEN IT FOCUS AT THE ORIGIN AND BECOME AN OUTGOING PACKET AT $t \rightarrow +\infty$ WITH A PHASE DIFFERENCE OF $\pi\ell$

FINALLY, PLANE WAVES

WE ARE NOW ABLE TO EXPAND THE PLANE WAVE $\langle \vec{r} | \vec{k} \rangle = \frac{e^{i\vec{k}\cdot\vec{r}}}{(2\pi)^{3/2}}$

IN TERMS OF THE FREE SPHERICAL WAVES $\psi_{k\ell m}^{(in)}$

THE RESULT IS $\langle \vec{r} | \vec{k} \rangle = \sum_{\ell} \frac{2\ell+1}{4\pi} \frac{1}{\sqrt{4\pi k}} c_{\ell} j_{\ell}(k r) P_{\ell}(\hat{k} \cdot \hat{r})$

IF WE CHOOSE $\vec{k} = k \hat{z}$ WITH $c_{\ell} = \frac{i^{\ell}}{k} \sqrt{\frac{2\ell+1}{\pi}}$

LEGENDRE POLYNOMIAL
 $= \sqrt{\frac{4\pi}{2\ell+1}} Y_{\ell 0}(\theta, \phi)$
 $\underbrace{\hspace{10em}}_{Y_{\ell 0}(\theta)}$

\Rightarrow THERE WILL BE NO DEPENDENCE ON ϕ

$\langle \vec{r} | 0, 0, k \rangle = \frac{e^{ikz}}{(2\pi)^{3/2}}$ (NOTICE $L_z |k\hat{z}\rangle = 0$ BECAUSE $L_z = x p_y - y p_x$)

AND $e^{ikz} = \sum_{\ell=0}^{\infty} i^{\ell} (2\ell+1) j_{\ell}(k r) P_{\ell}(\cos\theta)$

$$e^{ikz} = \sum_{\ell=0}^{\infty} i^{\ell} \sqrt{4\pi(2\ell+1)} j_{\ell}(k r) Y_{\ell 0}(\theta)$$

WE HAVE FINALLY SOLVED THE $V=0$ CASE

(7)

$$\psi^{(in)} = \frac{e^{ikz}}{(2a)^{1/2}} = \sum_{l=0}^{\infty} i^l \sqrt{\frac{2l+1}{2\pi^2}} j_l(kr) Y_{l,0}(\theta)$$

PARTIAL WAVES IN THE $V(r) \neq 0$ POTENTIAL

$$\psi_{k\ell m} = R_{k\ell}(r) Y_{\ell m}(\theta, \phi) = \frac{1}{r} u_{k\ell}(r) Y_{\ell m}(\theta, \phi)$$

$$\Rightarrow H\psi = E\psi \rightarrow \left(-\frac{\hbar^2}{2m} \frac{d^2}{dr^2} + \frac{l(l+1)\hbar^2}{2mr^2} + V(r) \right) u_{k\ell}(r) = \frac{\hbar^2 k^2}{2m} u_{k\ell}(r)$$

CONSTRAINED TO THE BOUNDARY CONDITION $u_{k\ell}(0) = 0$

FOR $r > a \Rightarrow V(r)$ IS NEGLIGIBLE

\rightarrow SOLUTIONS ARE $j_l(kr)$ AND $n_l(kr)$

WE CAN THEN TAKE THE $\lambda \rightarrow \infty$ LIMIT AND SIMPLIFY EVEN FURTHER

$$\left[-\frac{\hbar^2}{2m} \frac{d^2}{dr^2} u_{k\ell} = \frac{\hbar^2 k^2}{2m} u_{k\ell} \right]$$

$$\Rightarrow u_{k\ell} = A_l e^{ikr} + B_l e^{-ikr}$$

WHICH ARE THE BESSEL AND NEUMANN FOR $\lambda \rightarrow \infty$

A AND B CANNOT BE ARBITRARY

SINCE THE INCOMING WAVE IS TOTALLY REFLECTED INTO

THE OUTGOING WAVE (NO TRANSMISSION FOR $r < 0$) $\Rightarrow |A_l| = |B_l|$

$$\Rightarrow u_{k\ell} = C_\ell \sin\left(kr - \frac{\pi}{2}l + \delta_\ell\right)$$

\downarrow ($\delta_\ell \in \mathbb{R}$)

PHASE SHIFT DUE TO $V(r)$

$$\Rightarrow \psi_{k\ell m} \propto Y_{\ell m}(\theta, \phi) \frac{e^{i(kr - \frac{\pi}{2}l + \delta_\ell)} - e^{-i(kr - \frac{\pi}{2}l + \delta_\ell)}}{2} \propto Y_{\ell m}(\theta, \phi) \frac{e^{ikr} - e^{-i(kr - 2l\pi - 2\delta_\ell)}}{2}$$

→ COMPARING WITH THE ASYMPTOTIC BEHAVIOR OF THE FREE SPHERICAL WAVE, THIS NEW ONE CAN BE THOUGHT AS AN INCOMING SPHERICAL PACKET THAT FOCUSED ON THE ORIGIN AND IS TOTALLY REFLECTED AS AN OUTGOING WAVE. THE NET EFFECT OF THE POTENTIAL WAS THE ACCUMULATION OF AN EXTRA PHASE $2\delta_l$ RELATIVE TO THE $V=0$ CASE

- CONNECTION WITH S MATRIX → SEE PAGE 18.2

RELATION WITH THE SCATTERING AMPLITUDE

$$\psi_k = \frac{1}{(2\pi)^{3/2}} \left[e^{ikz} + f_k(\theta) \frac{e^{ikr}}{r} \right]$$

↓
RECALL $f_k(\theta, \phi) = f_k(\theta)$ FOR $V(\vec{x}) = V(r)$

WE HAVE FOUND THAT

$$\psi_k = \sum_{l=0}^{\infty} \sum_m c_l \underbrace{\sqrt{\frac{2l+1}{\pi}}}_{c_l} \frac{\sin(kr - \frac{l\pi}{2} + \delta_l)}{kr} Y_{lm}(\theta, \phi) \delta_{m,0}$$

ONLY $m=0$ BECAUSE THE INCOMING WAVE HAS $l_z=0$

$$= \frac{1}{(2\pi)^{3/2}} \left[\sum_{l=0}^{\infty} i^l \sqrt{4\pi(2l+1)} j_l(kr) Y_{l0}(\theta) + f_l(\theta) \frac{e^{ikr}}{r} \right]$$

↓
 $\approx \sin(kr - \frac{l\pi}{2}) / kr$

NOW, WE DO THE MATCHING:

• IDENTIFYING THE TERM $\propto e^{-i(kr - \frac{l\pi}{2})}$

$$\Rightarrow \frac{1}{(2\pi)^{3/2}} i^l \frac{\sqrt{4\pi(2l+1)} Y_{l0}}{2iKr} = -c_l \frac{Y_{l0}}{2iKr} e^{-i\delta_l}$$

$$\Rightarrow c_l = \frac{e^{i\delta_l}}{(2\pi)^{3/2}} i^l \sqrt{4\pi(2l+1)}$$

IT IS INTERESTING TO NOTICE THE CONNECTION WITH THE S (SCATTERING) MATRIX OPERATOR DEFINED BY

$$S = \lim_{\substack{t_f \rightarrow +\infty \\ t_i \rightarrow -\infty}} U(t_f, t_i)$$

where $U(t_f, t_i)$ IS THE ~~SCATTERING~~ TIME EVOLUTION OPERATOR $= e^{-i \frac{H(t_f - t_i)}{\hbar}}$

SINCE THE ONLY EFFECT OF SCATTERING IS A SIMPLE PHASE SHIFT (THIS SHOULD BE THE CASE BECAUSE S IS UNITARY AND CONSERVES ENERGY \Rightarrow THE EIGENVALUES OF S ARE SIMPLE PHASES $e^{i\alpha}$), THEN

WE CONCLUDE THAT THE MATRIX ELEMENT OF S ~~IS~~ WITH RESPECT TO THE WAVE OF ANGULAR MOMENTUM l IS $S_l = e^{2i\delta_l}$

NOTE THAT L^2 AND L_z ~~ARE~~ COMMUTE WITH H , THUS

THE S-MATRIX IS DIAGONAL IN THE ANGULAR MOMENTUM BASIS. THIS IS NOT THE CASE IN THE MOMENTUM BASIS SINCE \vec{p} IS NOT CONSERVED

WE CAN THEREFORE DEFINE

$$f(\theta) = \sum_{l=0}^{\infty} (2l+1) f_l(k) P_l(\cos\theta)$$

SEE PAGE 19

$$\Rightarrow f_l(k) = \frac{e^{i\delta_l} \sin \delta_l}{k} = \frac{e^{2i\delta_l} - 1}{2ik} \Rightarrow S_l = 1 + 2ik f_l(k)$$

IDENTIFYING THE TERM $\propto e^{ikr}$

$$\Rightarrow \sum_l c_l \frac{e^{i\delta_l}}{2ikr} e^{-i\delta_l} Y_{l,0} = \frac{1}{(2\alpha)^{3/2}} \left[\sum_l i^l \frac{\sqrt{4\pi(2l+1)}}{2ikr} e^{-i\delta_l} Y_{l,0} + \frac{f(\theta)}{r} \right]$$

$$\downarrow \frac{e^{i\delta_l}}{(2\alpha)^{3/2}} i^l \sqrt{4\pi(2l+1)}$$

$$\Rightarrow \sum_{l=0}^{\infty} \frac{i^l \sqrt{4\pi(2l+1)}}{2ik} \left(\frac{1}{k} - e^{2i\delta_l} \right) Y_{l,0} + f(\theta) = 0$$

$$\Rightarrow f(\theta) = \sum_{l=0}^{\infty} \frac{\sqrt{4\pi(2l+1)}}{k} \sin \delta_l e^{i\delta_l} Y_{l,0}(\theta)$$

RECALLING THAT $P_l(\cos\theta) = \sqrt{\frac{4\pi}{2l+1}} Y_{l,0}(\theta)$

$$\Rightarrow f(\theta) = \frac{1}{k} \sum_{l=0}^{\infty} (2l+1) e^{i\delta_l} \sin \delta_l P_l(\cos\theta)$$

~~TOTAL~~ CROSS-SECTION: $\sigma(\theta) = \frac{1}{k^2} \left| \sum_{l=0}^{\infty} (2l+1) e^{i\delta_l} \sin \delta_l P_l(\cos\theta) \right|^2$
 $= k^{-2} \sum_{l,l'} (2l+1)(2l'+1) e^{i(\delta_l - \delta_{l'})} \sin \delta_l \sin \delta_{l'} P_l(\cos\theta) P_{l'}(\cos\theta)$

TOTAL CROSS-SECTION

$$\sigma_{\text{tot}} = \int \sigma(\theta) d\Omega = \frac{2\pi}{k^2} \sum_{l,l'} (2l+1)(2l'+1) e^{i(\delta_l - \delta_{l'})} \sin \delta_l \sin \delta_{l'} \underbrace{\int P_l(\cos\theta) P_{l'}(\cos\theta) d\cos\theta}_{\frac{2}{2l+1} \delta_{l,l'}}$$

$$\Rightarrow \sigma_{\text{tot}} = \frac{4\pi}{k^2} \sum_{l=0}^{\infty} (2l+1) \sin^2 \delta_l$$

NOTICE THIS RESULT SATISFIES THE OPTICAL THEOREM

(*) $\text{Im } f(0) = \frac{k}{4\pi} \sigma_{\text{tot}}$

PROOF: $\text{Im } f(0) = \frac{1}{k} \sum_{l=0}^{\infty} (2l+1) \sin^2 \delta_l P_l(1)$

SINCE $P_l(x) = 2^l \sum_{k=0}^l x^k \binom{l}{k} \left(\frac{x+k-1}{2}\right)$

$\Rightarrow P_l(1) = 2^l \sum_{k=0}^l \binom{l}{k} \left(\frac{l+k-1}{2}\right) = 2^l \times 2^{-l} = 1$

ALTERNATIVELY, WE COULD USE THAT

$(l+1)P_{l+1}(x) = (2l+1)xP_l(x) - lP_{l-1}(x)$

SINCE $P_0(x) = 1$ AND $P_1(x) = x$

\Rightarrow FOR $l=1 \rightarrow 2P_2(1) = 3 - 1 = 2 \rightarrow P_2(1) = 1$

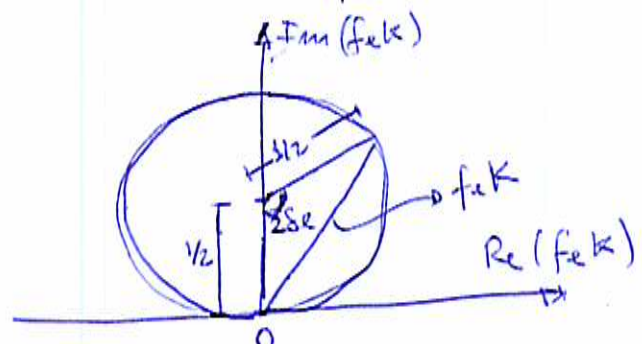
INDUCTIVE HYPOTHESIS $P_l(1) = 1 = P_{l-1}(1) \rightarrow (l+1)P_{l+1}(1) = 2l+1 - l = l+1 \Rightarrow P_{l+1}(1) = 1$

$\Rightarrow \text{Im } f(0) = \frac{1}{k} \sum_{l=0}^{\infty} (2l+1) \sin^2 \delta_l = \frac{k}{4\pi} \left(\frac{4\pi}{k^2} \sum_{l=0}^{\infty} (2l+1) \sin^2 \delta_l \right) = \frac{k}{4\pi} \sigma_{\text{tot}}$

PARTIAL CROSS-SECTION $\sigma_{\text{tot}} = \sum_l \sigma_{l,\text{tot}} = 0 \quad \sigma_{l,\text{tot}} = \left(\frac{4\pi}{k^2}\right)^2 (2l+1) \sin^2 \delta_l$

where $\tilde{\lambda} = \frac{1}{k}$. NOTICE THE MAXIMUM PARTIAL CROSS SECTION IS ACHIEVED WHEN $\delta_l = \pm \frac{\pi}{2}$, IMPLYING A MAXIMAL $|f_e(k)|$

RECALL $f_e(k) = e^{i\delta_l} \sin \delta_l = \frac{1}{2i} (e^{2i\delta_l} - 1) = \frac{\cos(2\delta_l) - 1 + i \sin(2\delta_l)}{2i} = \frac{1}{2} (\sin(2\delta_l) + i(1 - \cos(2\delta_l)))$



NOW, WE NEED TO COMPUTE THE PHASE SHIFTS
(OTHERWISE, WE HAVE DONE NOTHING)

THIS IS NOT A SIMPLE TASK IN GENERAL. THUS, WE WILL FOCUS IN ONE EXAMPLE: HARD-SPHERE POTENTIAL

$$V(r) = \begin{cases} \infty, & r < a \\ 0, & r > a \end{cases}$$

AS DISCUSSED ON PAGE (17), $R_{kl}(r) = A_l j_l(kr) + B_l n_l(kr)$
FOR $r > a$, AND $R(r) = 0$, $r \leq a$

FROM THE CONTINUITY OF $R_{kl} = 0$ $A_l j_l(ka) + B_l n_l(ka) = 0$

THE ASYMPTOTIC BEHAVIOR OF $R_{kl} \xrightarrow{r \rightarrow \infty} \frac{A_l \sin(kr - \frac{r\pi}{2}) + B_l \cos(kr - \frac{r\pi}{2})}{kr}$

CALLING $A_l = C_l \cos \delta_l$
 $B_l = -C_l \sin \delta_l \Rightarrow R_{kl} \rightarrow \frac{C_l}{kr} \sin(kr - \frac{r\pi}{2} + \delta_l)$

THEREFORE, $\tan \delta_l = -\frac{B_l}{A_l} = \frac{j_l(ka)}{n_l(ka)} = \tan \delta_l$

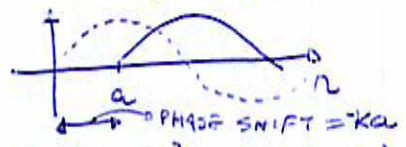
FOR SMALL ka \Rightarrow ~~$n_l(ka) \approx \frac{(2l-1)!!}{(ka)^{2l+1}} (1 + \frac{(ka)^2}{2(2l-1)} + \dots)$~~

$$\Rightarrow \tan \delta_l \approx -\frac{(ka)^{2l+1}}{2l+1} \left(\frac{1}{(2l-1)!!} \right)^2$$

$$j_l = \frac{(ka)^l}{(2l+1)!!} \left(1 - \frac{(ka)^2}{2(2l+3)} + O(ka)^4 \right)$$

$$n_l = -\frac{(2l-1)!!}{(ka)^{2l+1}} \left(1 + \frac{(ka)^2}{2(2l-1)} + O(ka)^4 \right)$$

$$\Rightarrow \delta_0 \approx -ka$$



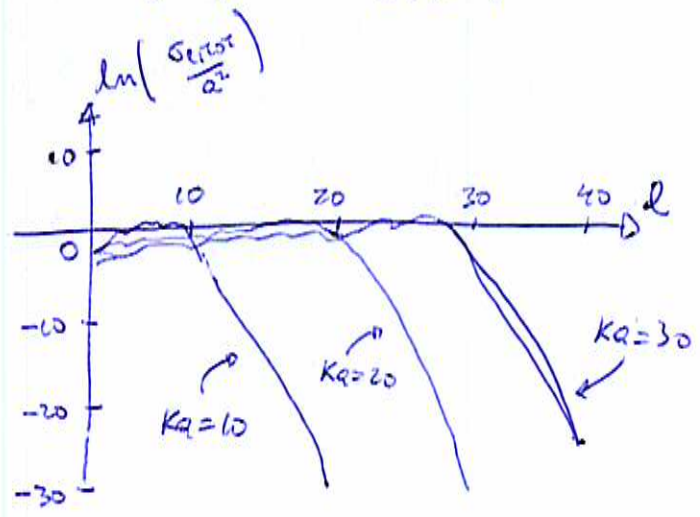
$$\Rightarrow \sigma_{tot} \approx \sigma_{l=0, tot} \approx \frac{4\pi a^2}{k^2} (ka)^2 = 4(\pi a^2)$$

$= 4 * (\text{CLASSICAL } \sigma_{tot})$

HOW MANY l 's WE HAVE TO KEEP?
APPROXIMATION IS GOOD UP TO

$l \sim ka$, i.e., WE HAVE TO CONSIDER $\sigma_{l, tot}$ UP TO $l \sim ka$

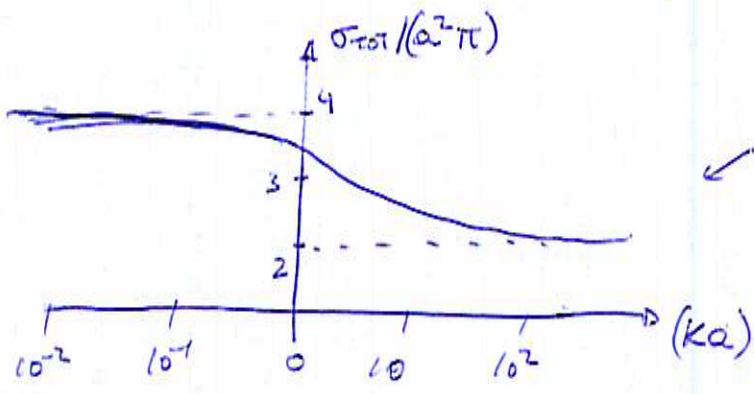
$$\sigma_{tot} \approx \sum_{l=0}^{ka} \sigma_{l, tot}$$



THIS RESULT, AGAIN, IS OKAY WITH THE

INTUITION THAT FOR $b_e = \frac{L}{P} = \frac{h/\lambda(1+n)}{hk} \sim \frac{L}{k}$ GREATER

THAN THE RADIUS $a \Rightarrow \delta_e \approx 0$ AND NOTHING HAPPENS FOR THOSE PARTICLE WAVES ($\lambda > \lambda_{max} \sim ka$)



FOR LARGE ka , IT IS NOT CONVERGING TO πa^2 , BUT INSTEAD TO $2\pi a^2$

FOR LARGE ka $\rightarrow j_e(ka) \approx \frac{1}{ka} \sin(ka - \frac{1}{2}\pi)$

$n_e(ka) \approx \frac{1}{ka} \cos(ka - \frac{1}{2}\pi)$

$$\sin^2 \delta_e = \frac{\tan^2 \delta_e}{1 + \tan^2 \delta_e} = \frac{j_e^2}{j_e^2 + n_e^2} \approx \sin^2(ka - \frac{1}{2}\pi)$$

$$\Rightarrow \sin^2 \delta_e + \sin^2 \delta_{e+1} \approx \sin^2 \delta_e + \sin^2(\delta_e - \frac{\pi}{2}) = \sin^2 \delta_e + \cos^2 \delta_e = 1$$

$$\begin{aligned} \Rightarrow \sigma_{tot} &= \frac{4\pi}{k^2} \sum_{l=0}^{\infty} (2l+1) \sin^2 \delta_l = \frac{4\pi}{k^2} \left(\sin^2 \delta_0 + \sin^2 \delta_1 + 2\sin^2 \delta_1 + 2\sin^2 \delta_2 + 3\sin^2 \delta_2 + 3\sin^2 \delta_3 + 4\sin^2 \delta_3 + \dots + (ka-1)\sin^2 \delta_{ka-1} + ka \sin^2 \delta_{ka-1} + ka \sin^2 \delta_{ka} + (ka+1)\sin^2 \delta_{ka} \right) \\ &\approx \frac{4\pi}{k^2} \sum_{l=0}^{ka} (2l+1) \sin^2 \delta_l \approx \frac{4\pi}{k^2} \times \frac{1}{2} (ka+2)ka \end{aligned}$$

$\sigma_{tot} \approx 2\pi a^2$ NOT THE GEOMETRIC CROSS SECTION, WHY NOT?

$$f(\theta) = \frac{1}{k} \sum_{l=0}^{ka} (2l+1) e^{i\delta_l} \sin \delta_l P_l(\cos \theta) = \frac{1}{2ik} \sum_{l=0}^{ka} (2l+1) e^{2i\delta_l} P_l(\cos \theta) + \frac{i}{2k} \sum_{l=0}^{ka} (2l+1) P_l(\cos \theta)$$

= REFLECTION + TRANSMISSION

$$\sigma_{\text{REF, TOT}} \equiv \int |f_{\text{REF}}|^2 d\Omega \quad (23)$$

$$= \frac{1}{4k^2} \sum_{l=0}^{ka} (2l+1)^2 \underbrace{\int P_l^2(\cos\theta) d\Omega}_{\frac{4\pi}{2l+1}} = \frac{\pi}{k^2} \sum_{l=0}^{ka} (2l+1) \underbrace{(k+l)ka}_{\approx k^2 a^2}$$

$$\Rightarrow \boxed{\sigma_{\text{REF, TOT}} = \pi a^2}$$

ANALOGOUSLY, $\sigma_{\text{SHADOW, TOT}} \equiv \int |f_{\text{SHADOW}}|^2 d\Omega$

with $f_{\text{SHADOW}} \approx \frac{i}{2k} \sum_l (2l+1) P_l(\cos\theta)$

NOTICE THAT f_{SHADOW} IS PURELY IMAGINARY AND ADDS UP COHERENTLY FOR $\theta = 0$

\Rightarrow BRIGHT SPOT BEHIND THE POTENTIAL

FOR LARGE l AND SMALL θ

$P_l(\cos\theta) \approx J_0(l\theta)$ $\xrightarrow{\text{BESSEL FUNCTION}}$

$$\Rightarrow f_{\text{SHADOW}} \approx \frac{i}{2k} \sum_{l=0}^{ka} (2l+1) J_0(l\theta) \xrightarrow{\text{LARGE } ka} \frac{i}{2k} \frac{1}{\Delta l} \int_0^{ka} dl \, 2l J_0(l\theta), \Delta l = 1$$

USING THAT $b = \frac{a}{k} \rightarrow \frac{i}{k} \cdot k^2 \int_0^a db \, b J_0(k\theta b) = ik \times \frac{1}{(k\theta)^2} (1 - \cos(k\theta a))$

$$= \frac{ia}{\theta} \frac{1 - \cos(k\theta a)}{k\theta a} \approx \frac{ia}{\theta} J_1(k\theta a) \rightarrow \text{FRAUNHOFER DIFFRACTION PATTERN}$$

JUST AS $\sigma_{\text{REF}} \rightarrow \sigma_{\text{SHADOW}} = \frac{1}{4k^2} \sum_l (2l+1)^2 \int P_l^2(\cos\theta) d\Omega = \pi a^2$

FINALLY, THE INTERFERENCE ~~REFF~~ BETWEEN σ_{REF} AND σ_{SHAD}

VANISHES $\text{Re}(f_{\text{SHAD}}^* f_{\text{REF}}) \propto \text{Re}\left(\sum_l e^{2i\delta_l}\right) \rightarrow 0$

BECAUSE $2\delta_{l+1} = 2\delta_l - \pi \Rightarrow$ STRONG OSCILLATIONS AVERAGING TO ZERO

FINALLY, NOTICE THAT

(24)

$$\sigma_{\text{TOT}} = \frac{4\pi}{k} \text{Im}(f(0)) = \frac{4\pi}{k} \text{Im}(f_{\text{SHADOW}}(0)) \quad \text{BECAUSE}$$

$\text{Im } f_{\text{REF}}(0) \propto \text{Im} \left(\sum_e e^{2ikz} \right)$ WHICH STRONGLY OSCILLATES
AVERAGING TO ZERO.

$$\text{AND } \text{Im} \left(f_{\text{SHADOW}}(0) \right) = \frac{1}{2k} \sum_{z=0}^{kr} 2kz \approx \frac{1}{2k} \cdot 2 \frac{(ka)^2}{2}$$

$$\Rightarrow \sigma_{\text{TOT}} = \frac{4\pi}{k} \cdot \frac{1}{2} ka^2 = \underline{2\pi a^2 = \sigma_{\text{TOT}}}$$

